

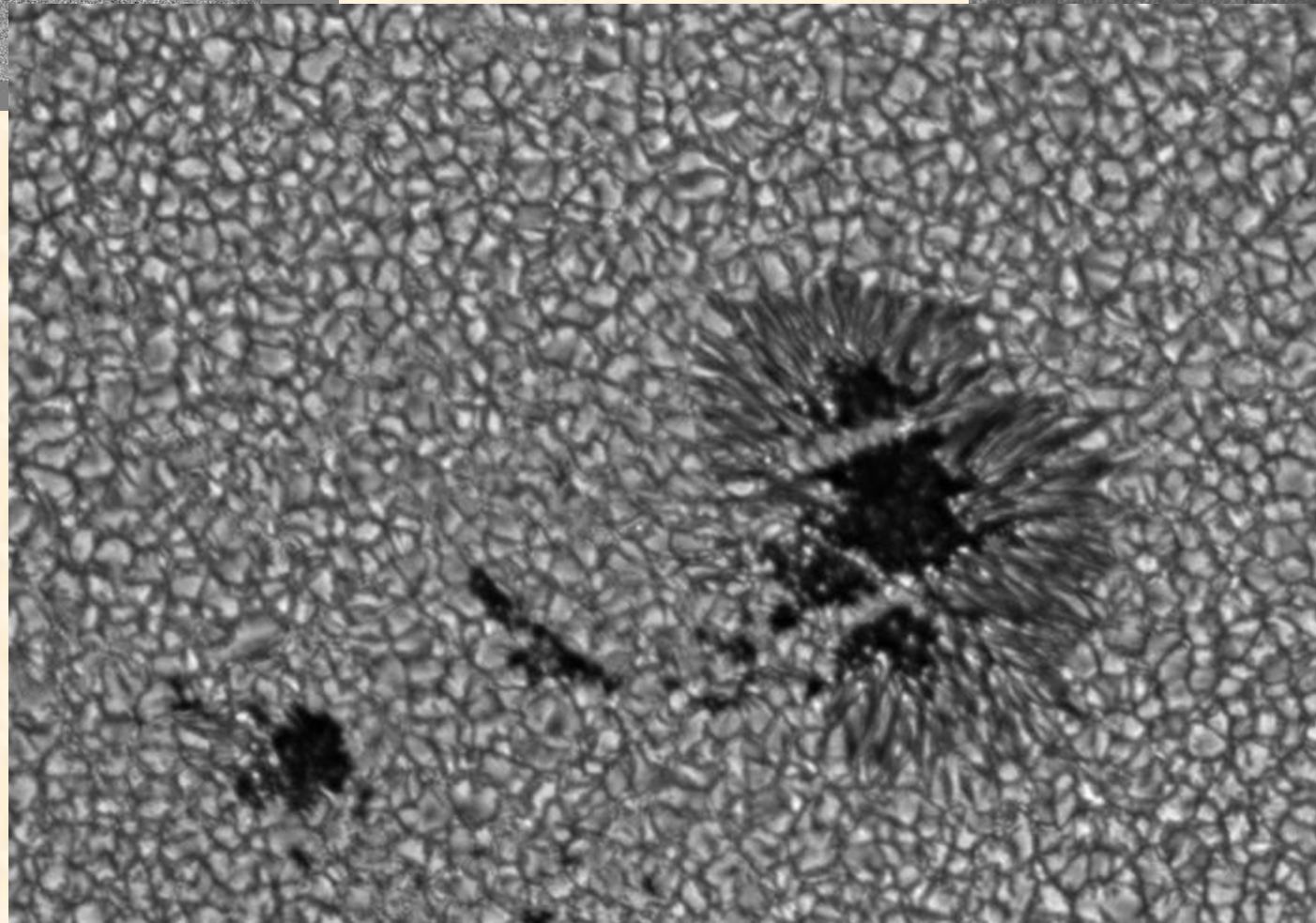
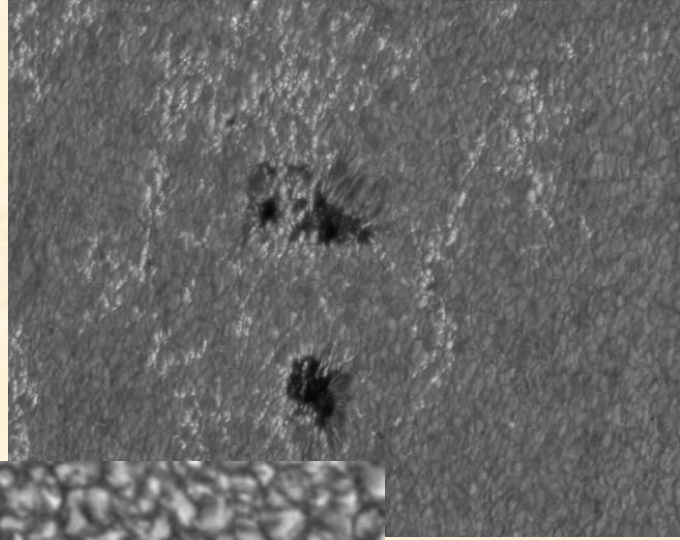
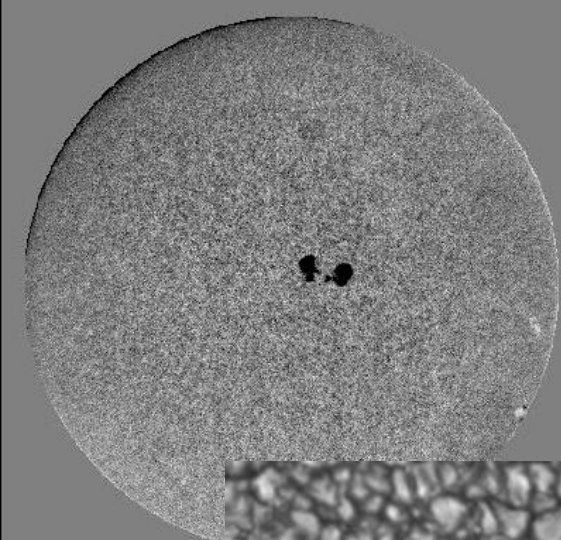
***The asymmetry of the solar
cycle: Analysis of the cycle
#24 forecast.***

Mircea V. Rusu

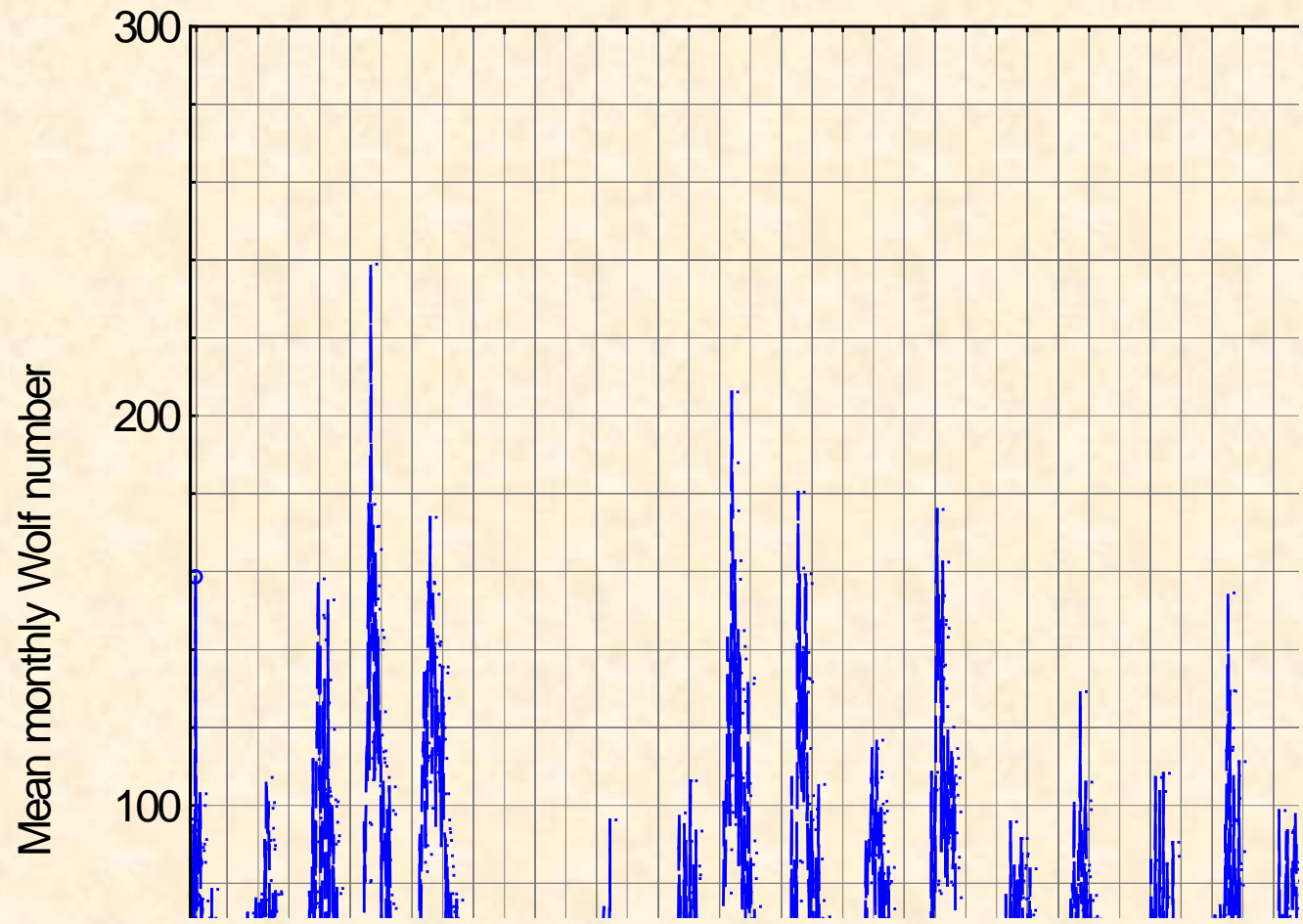
Physics Department, Bucharest University,
Bucharest, Romania

Introduction

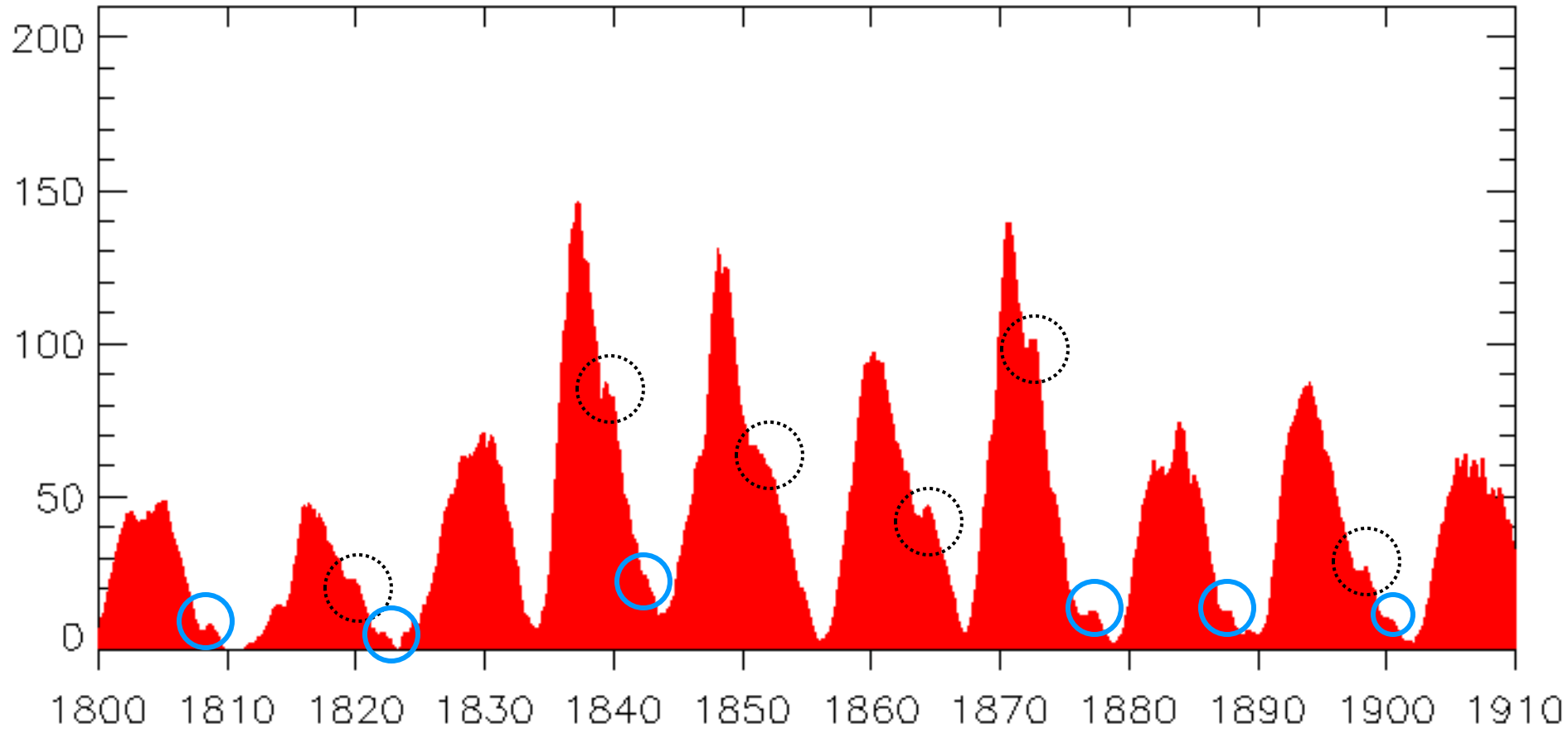
- Different numerical analysis models for fitting the shape of Wolf number evolution – not satisfactory
- Few attempts to base the forecast on a physical basis
- The paper presents the progress we made in analysis of the Wolf sunspot numbers with a special attention on the asymmetry exhibited for each cycle based on our previous findings:
 - Periodicity of approximately 11 years of maxima
 - Variability of the “amplitude” of these maxima
 - Asymmetry of the variation: a quick increase and a much slower decrease of the sunspot numbers (maxima present asymmetry)
 - Many of them the presence of two maxima



The pattern of the time series of wolf numbers for a log period

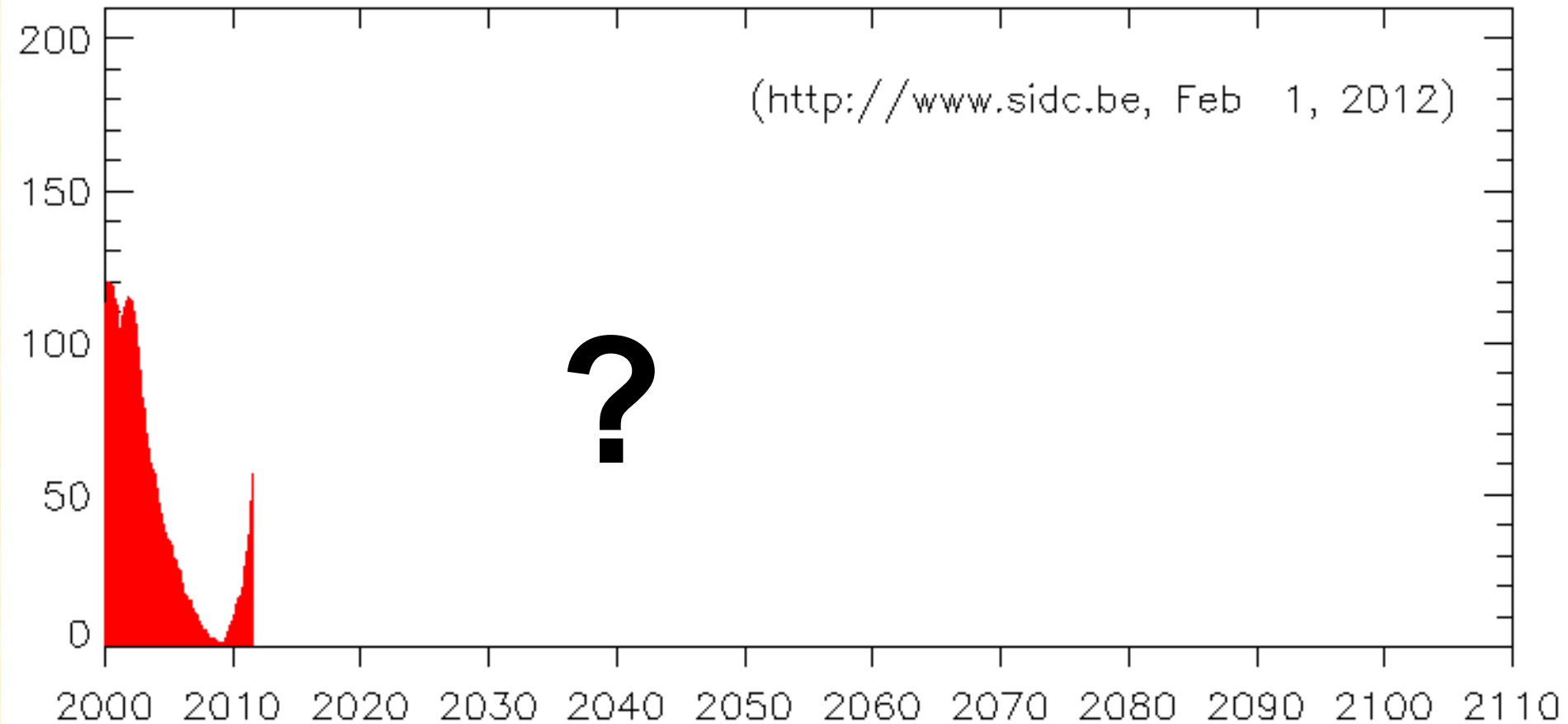


The pattern of the time series of wolf numbers for a log period

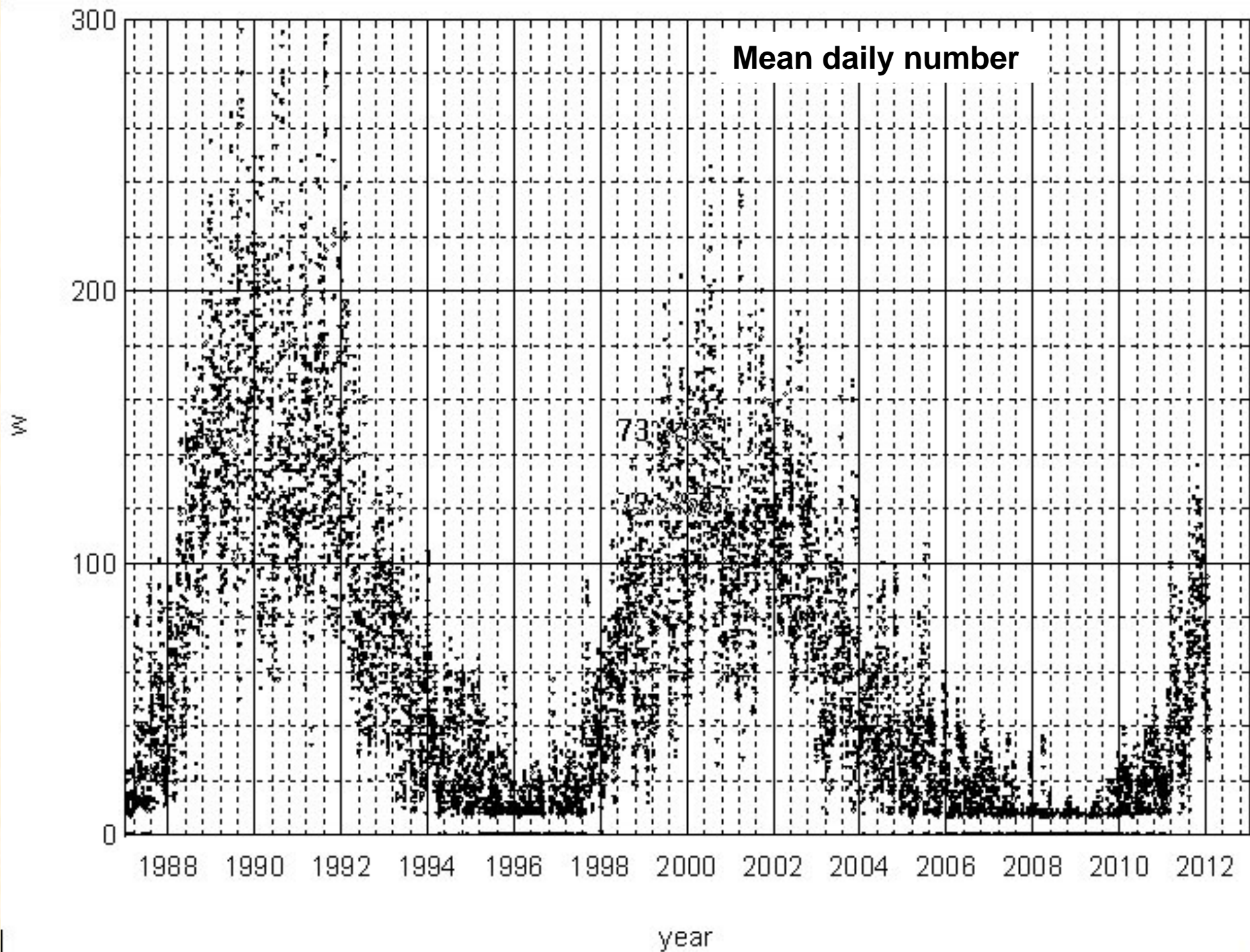


Mean monthly number

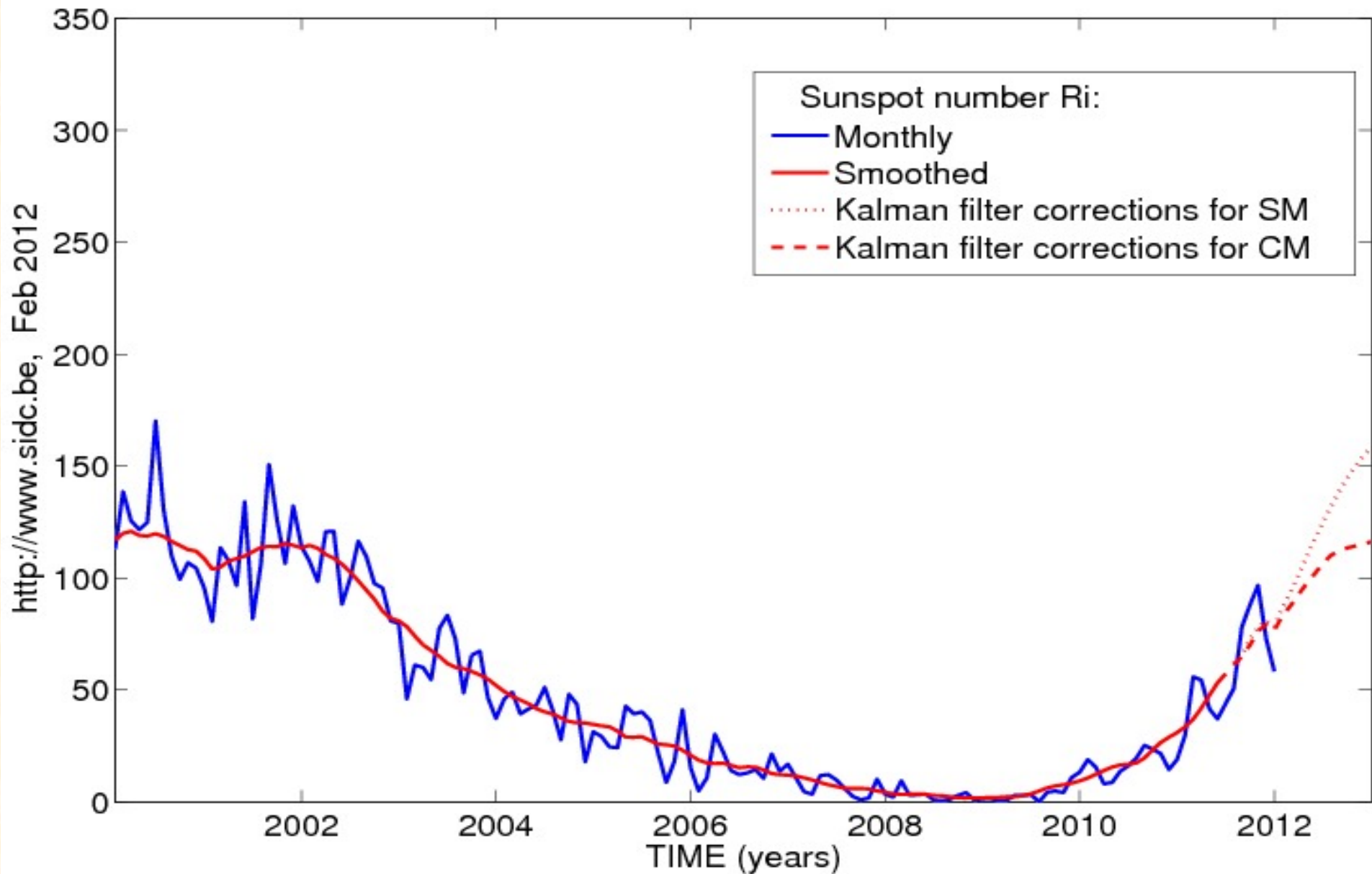
The pattern of the time series of wolf numbers for a log period



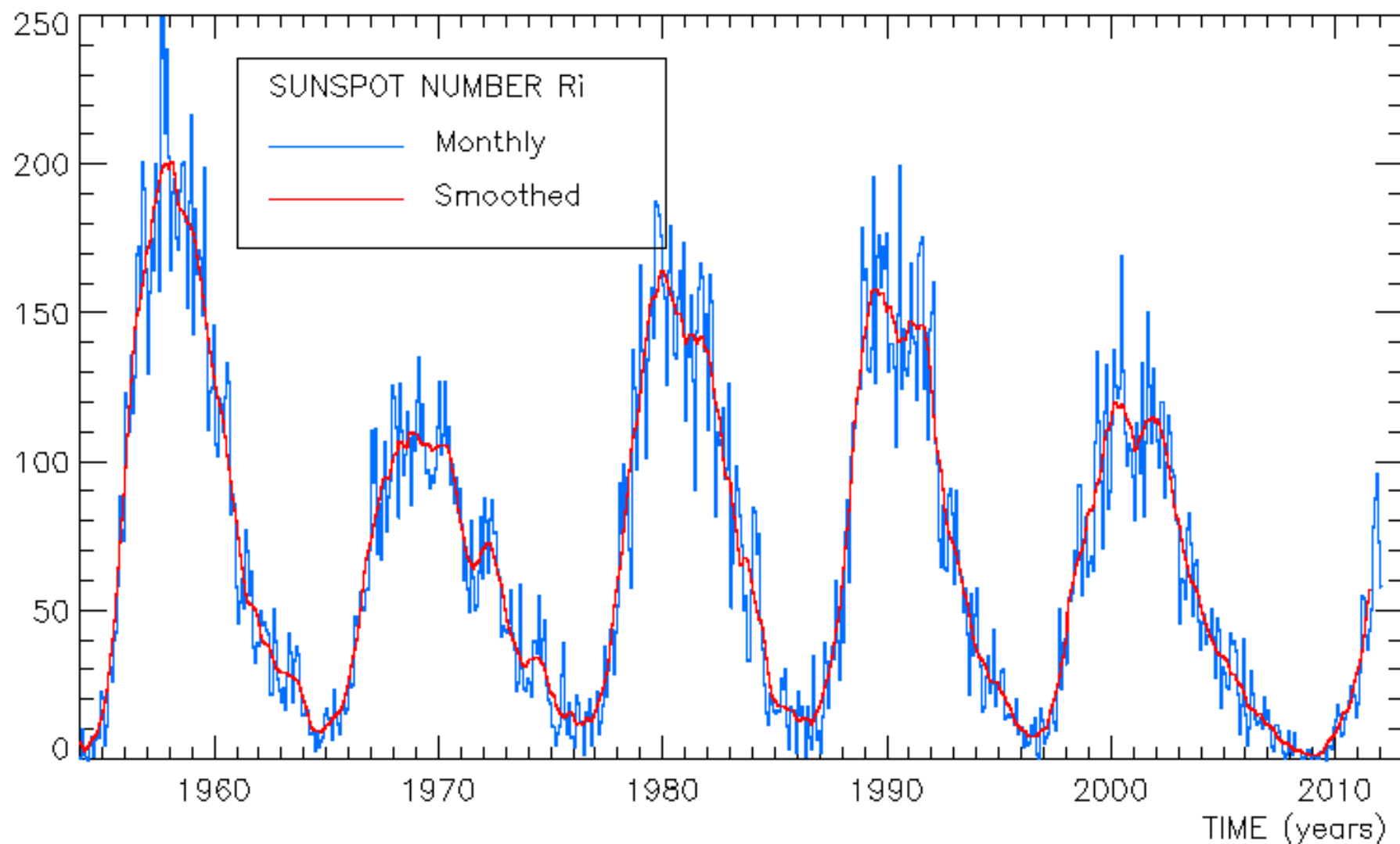
Mean monthly number



General models – smoothed

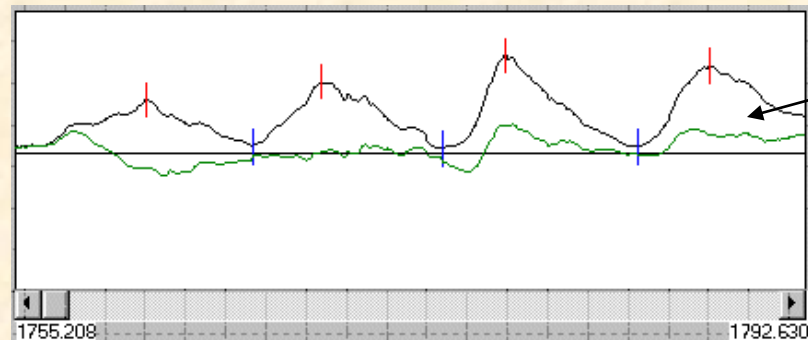
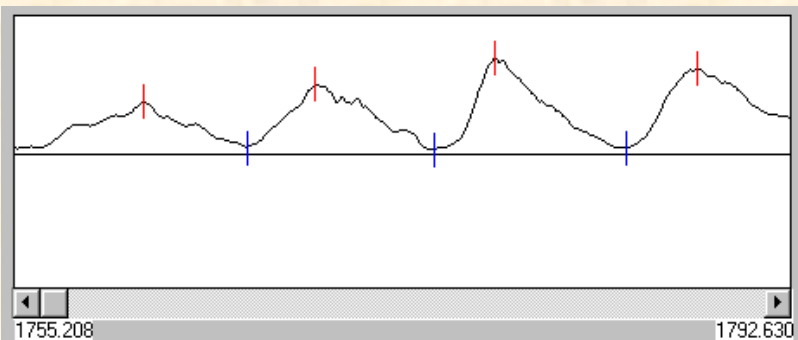
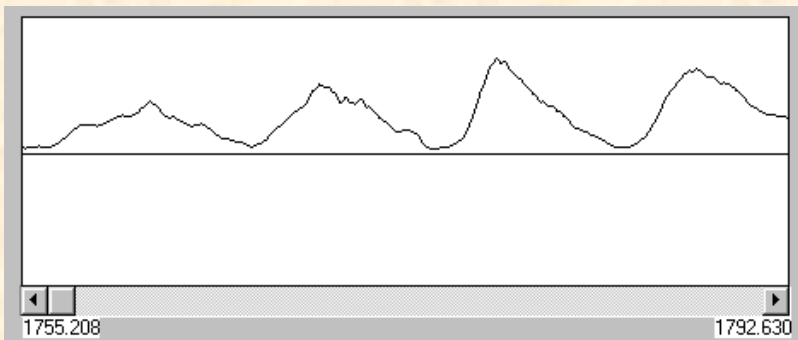


Pattern for the mean values – sliding mean

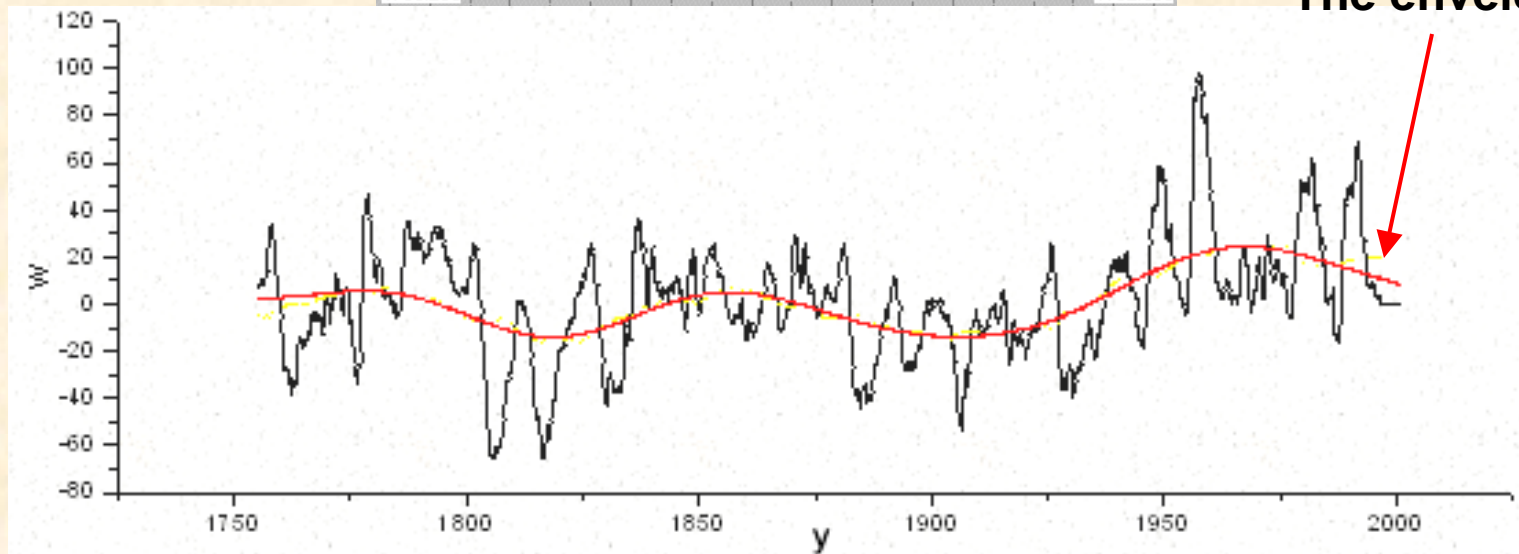


Our procedure for finding the minima & maxima of the wolf number series

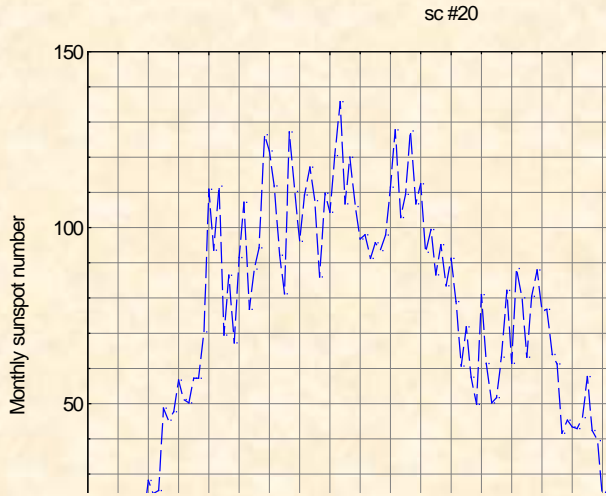
(O. Tesileanu, Z. Mouradian, M. Rusu – 2002)



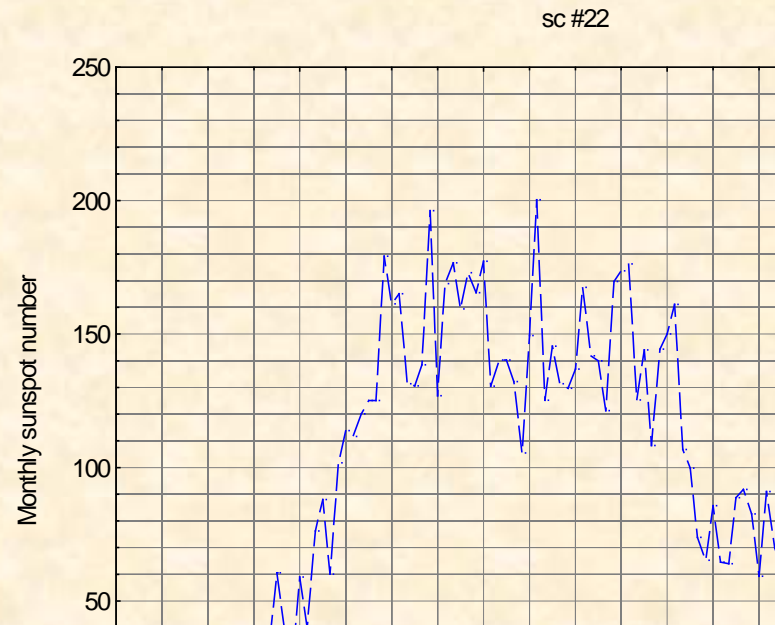
The difference



The envelope

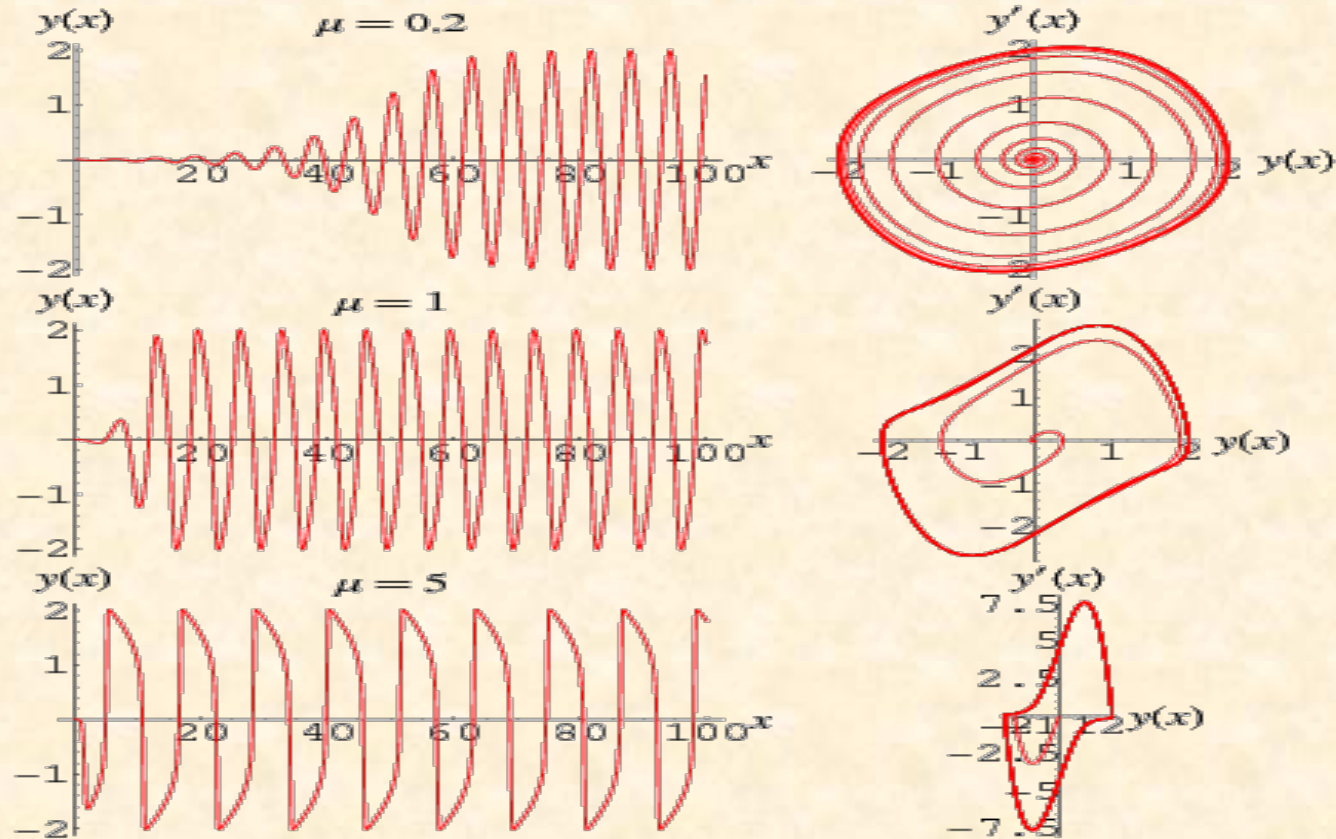


The asymmetry



Nonlinear van der Pol model

(J.M. Polygiannakis, X. Moussas, C.P. Sonett - 1996)

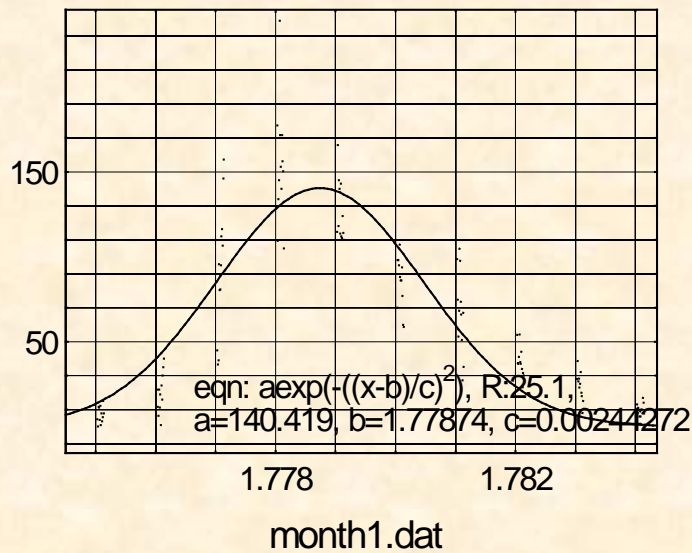


Time series and space phase plot for van der Pol oscillator

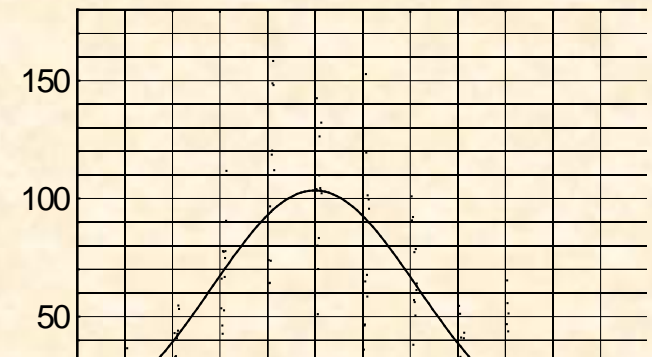
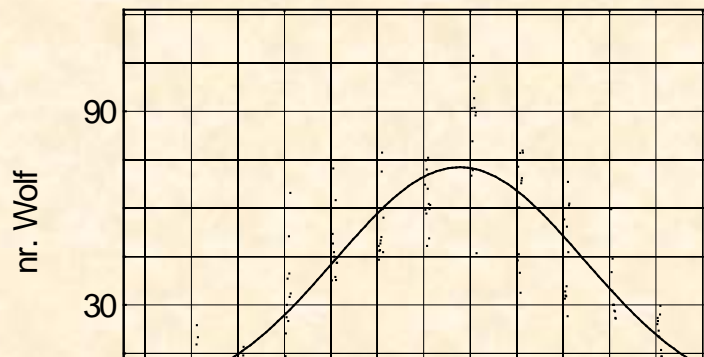
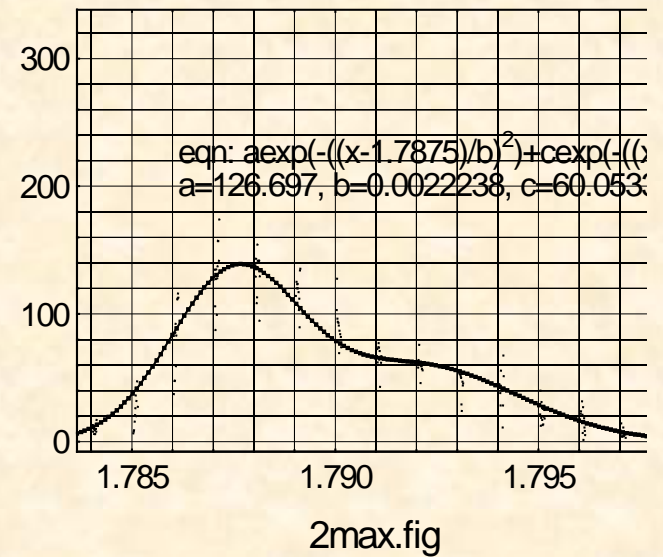
A detailed analysis of each cycle – samples

(M.V.Rusu – 2004 - 2006)

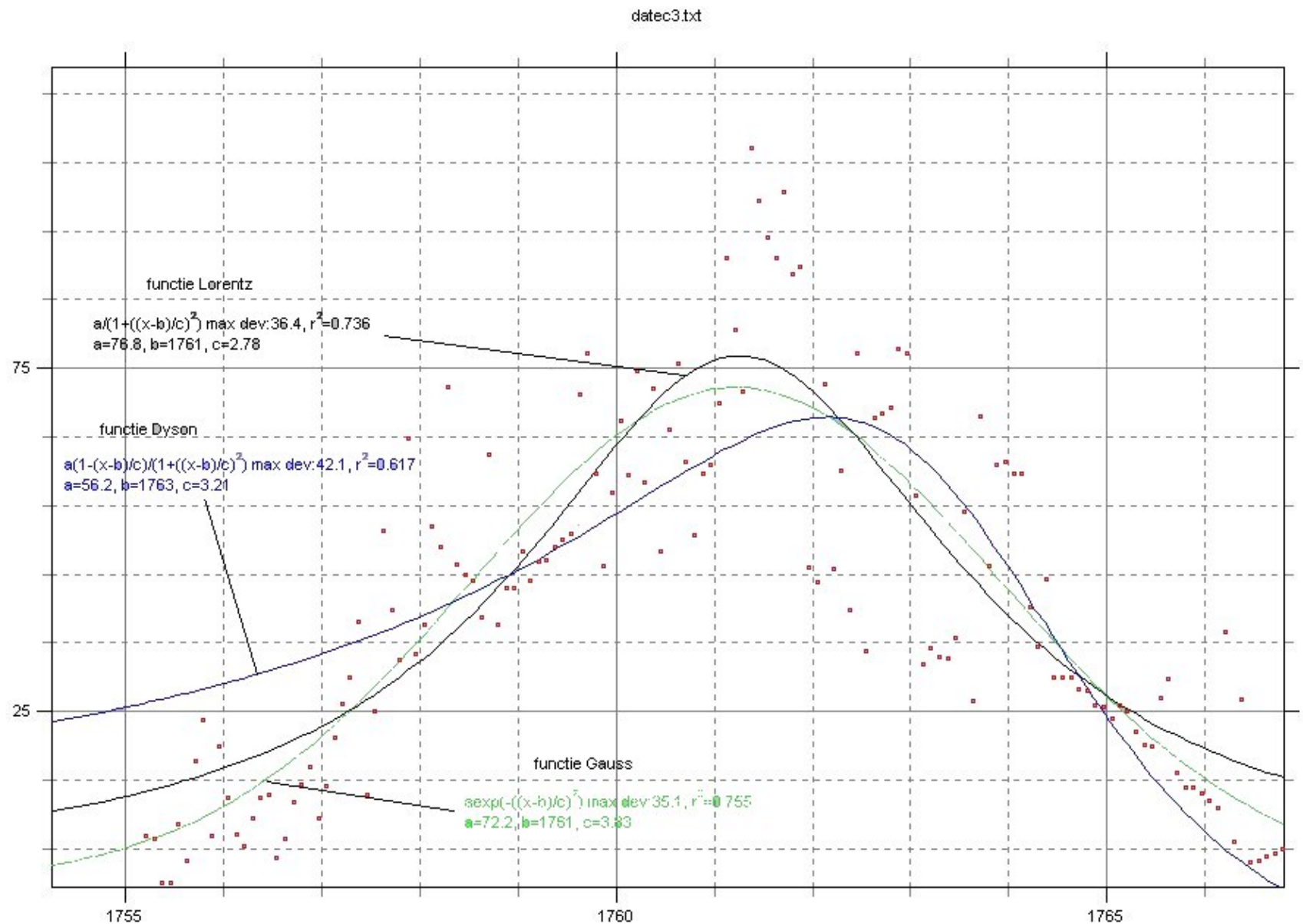
al 3-lea maxim



4max.fig

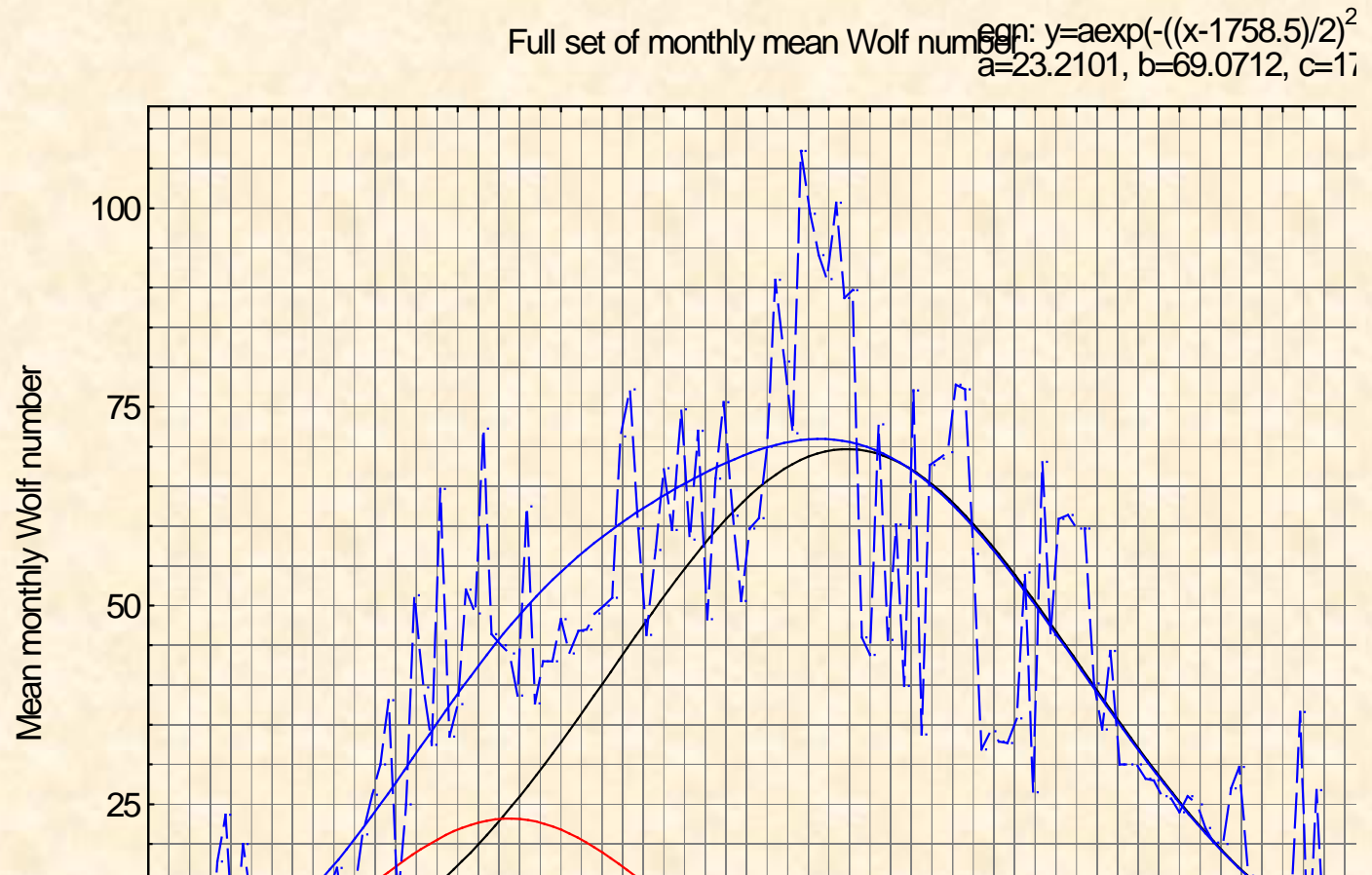


Search for the best fit symmetrical curve: **Gauss**, Lorenz, Dyson

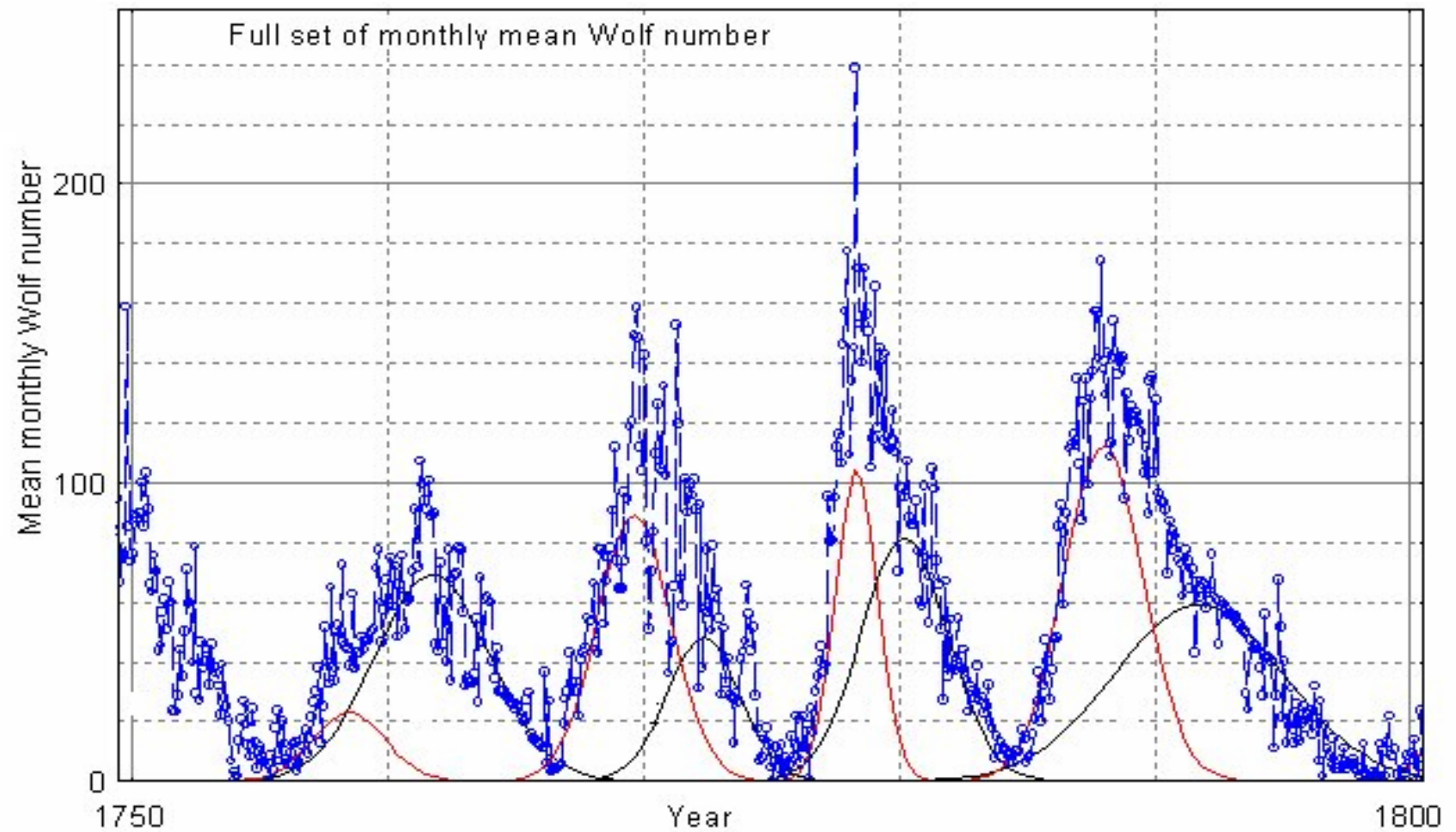


De-convolution

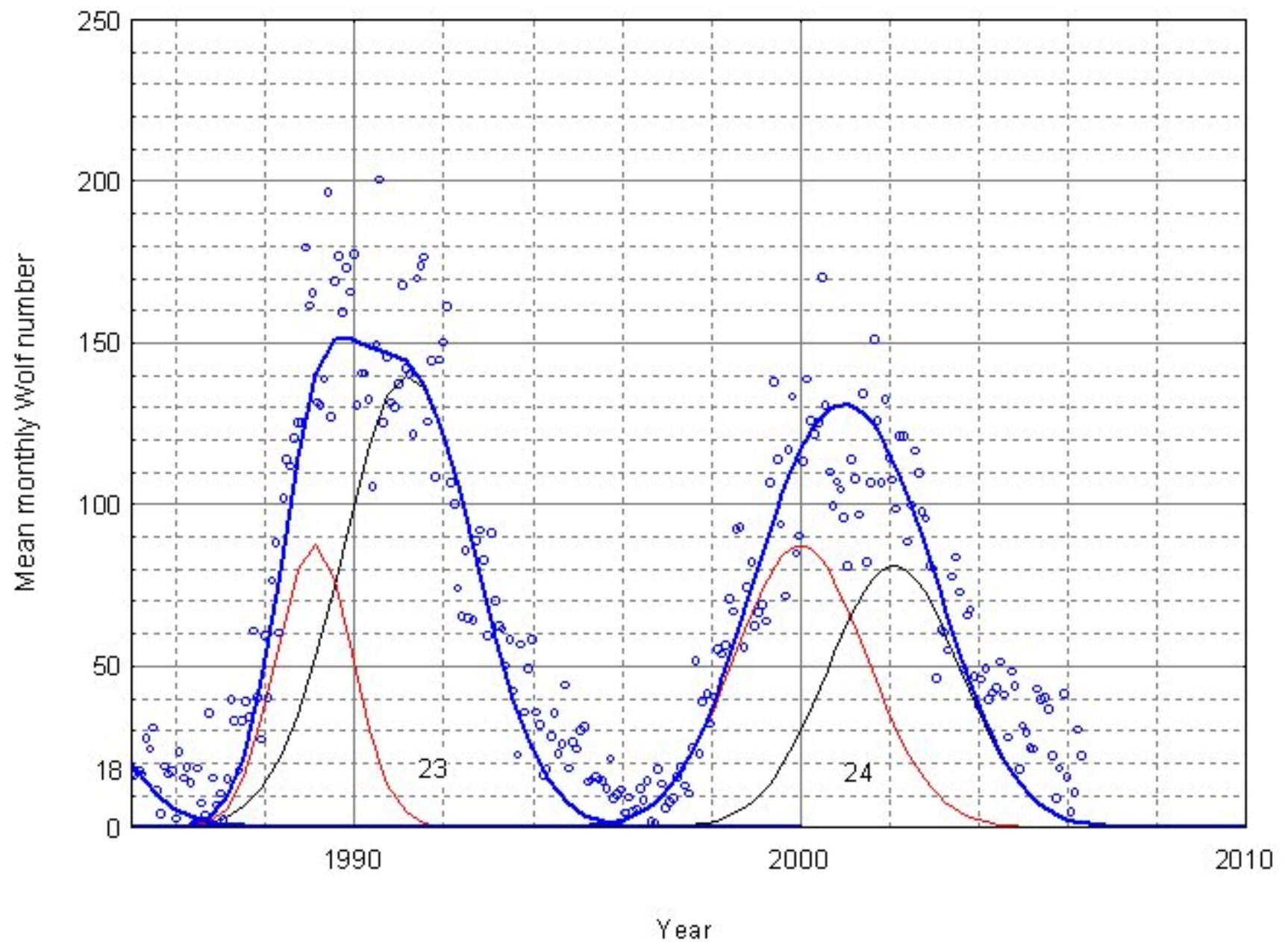
The idea: **the asymmetry is due to presence of superposition of two peaks**



De-convolution



Full set of monthly mean Wolf number



$$f(x) = a \cdot e^{-\left(\frac{x-b}{c}\right)^2} + d \cdot e^{-\left(\frac{x-f}{g}\right)^2}$$

Gaussian functions - describe the normal distributions (statistics)
 - solve heat equations and diffusion equations (math);

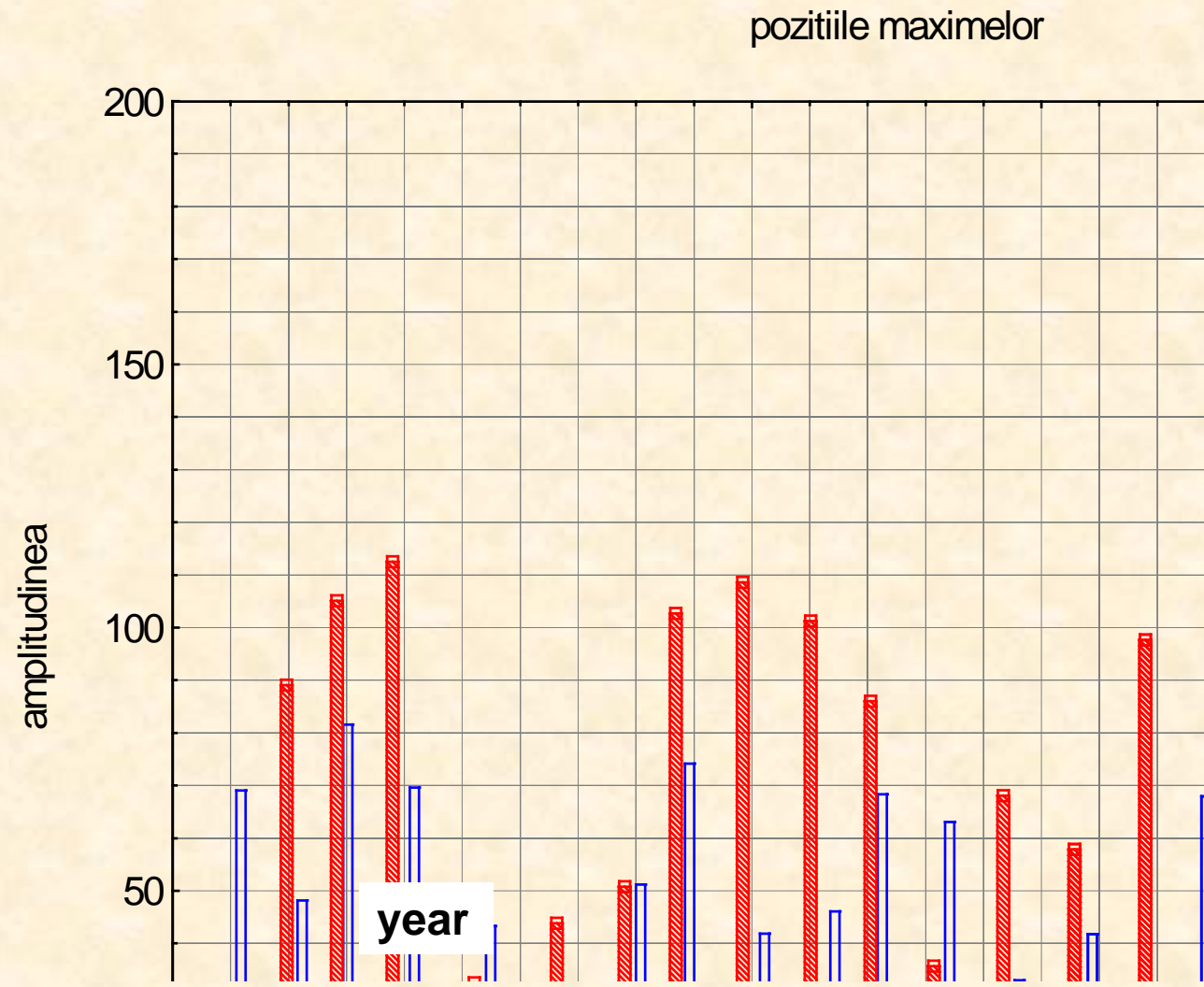
Mean $\mu = b$, Variance $\sigma^2 = c^2$: Area $a \cdot c \cdot \sqrt{2\pi}$

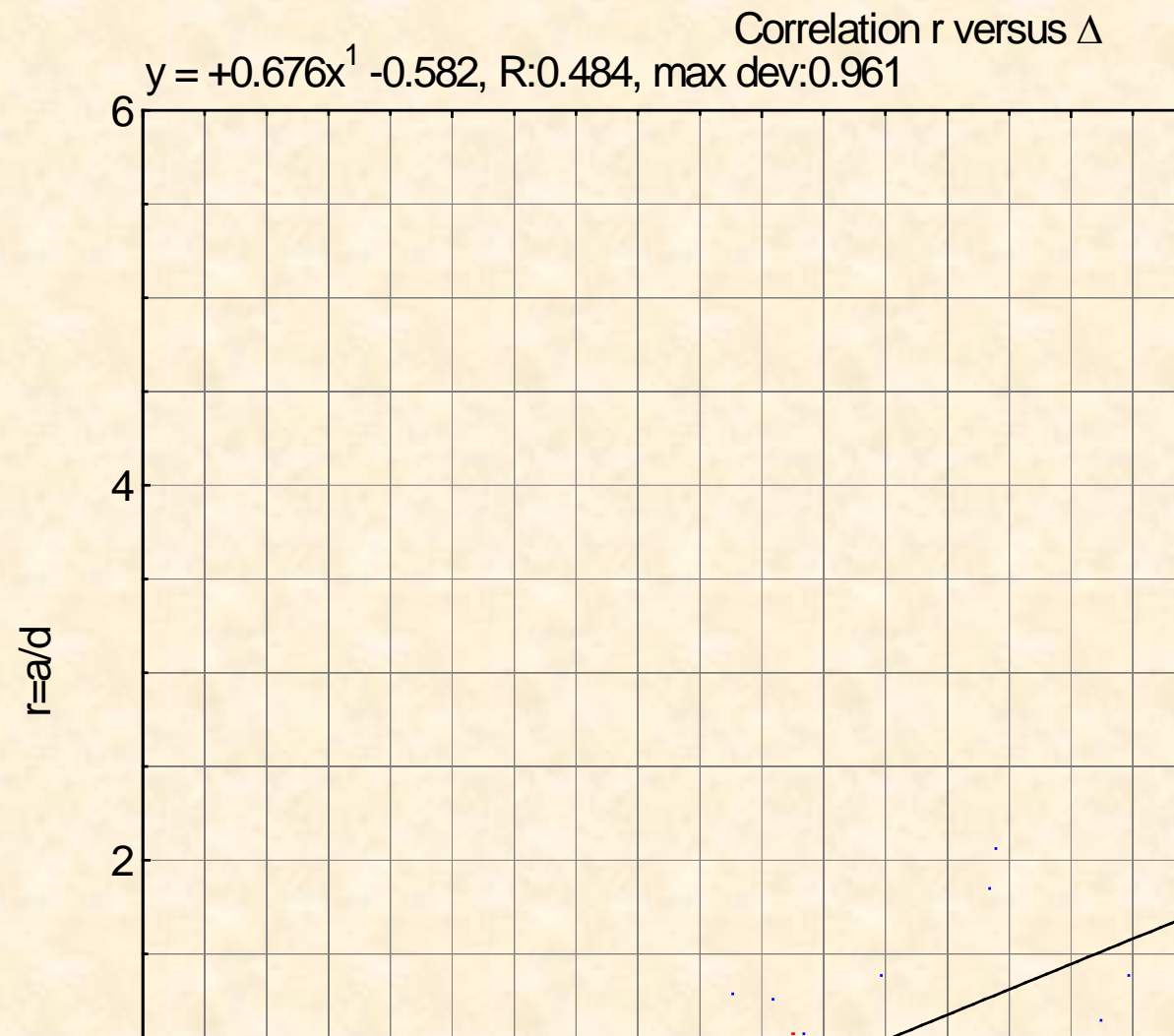
The product of two Gaussian functions is a Gaussian,

The convolution of two Gaussian functions is a Gaussian, with $\Delta = \sqrt{c^2 + g^2}$

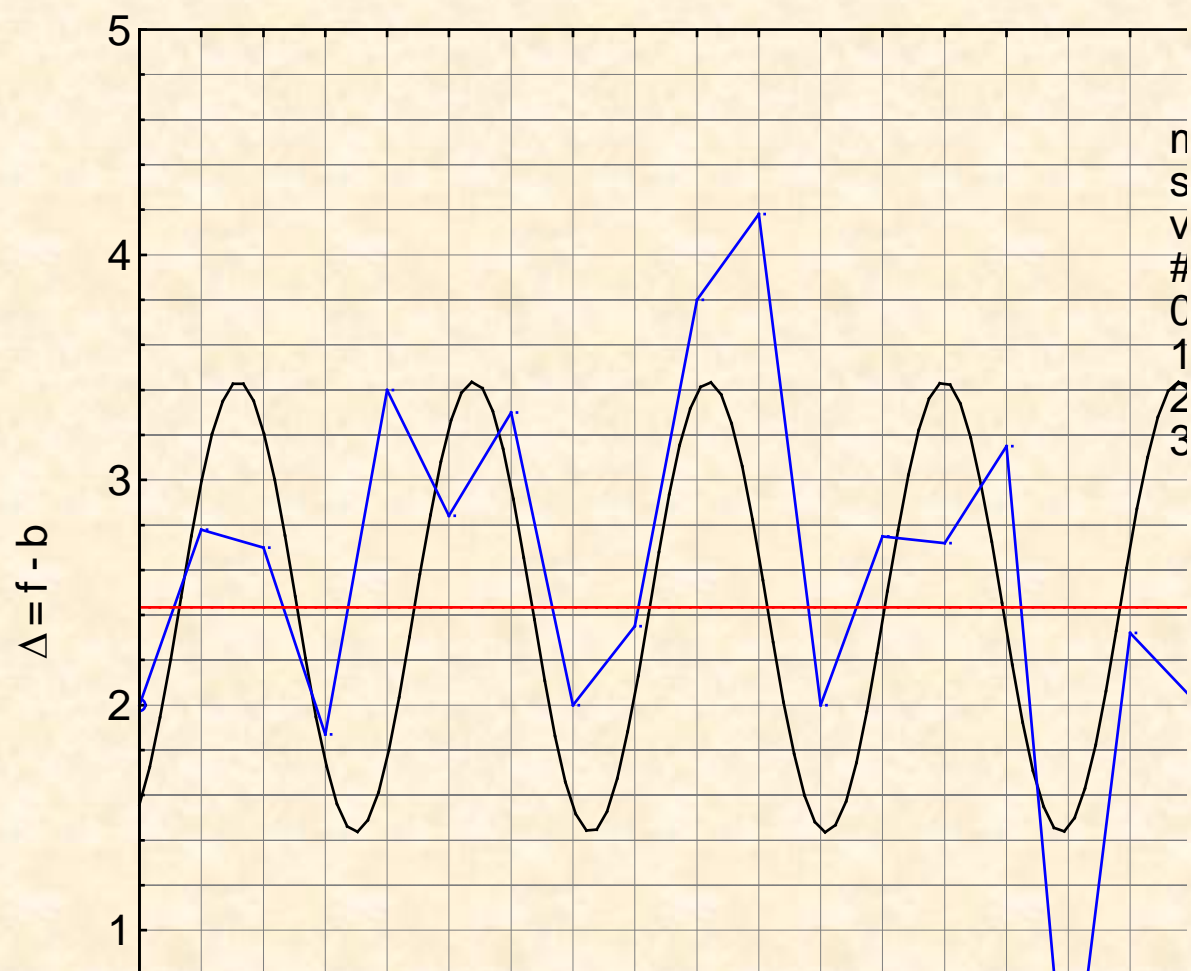
No.	A					B					Δ	R
Cycle.	b poz 1	a ampl 1	c width 1	a/c	a*c area A	f poz 2	d ampl 2	g width 2	d/g	d*g area B	f - b	a/d
1	1759.03	23.21	1.05	22.10	24.3705	1761.78	69.07	3.23	21.38	223.0961	2.75	0.336
2	1769.67	89.10	2.04	43.67	181.764	1772.37	48.19	2.20	21.90	106.018	2.70	1.849
3	1778.35	105.07	1.16	90.58	121.8812	1780.22	81.59	2.40	34.00	195.816	1.87	1.288
4	1788.03	112.51	2.20	51.14	247.522	1791.40	64.09	4.40	14.56	281.996	3.37	1.755
5	1802.05	36.54	1.44	25.37	52.6176	1804.89	45.34	2.03	22.33	92.0402	2.84	0.806
6	1816.20	48.00	2.09	22.97	100.32	1819.50	22.03	1.50	14.69	33.045	3.30	2.179
7	1828.04	50.79	2.22	22.88	112.7538	1830.00	51.23	2.30	22.27	117.829	1.96	0.991
8	1836.88	102.71	1.43	71.82	146.8753	1839.23	74.16	2.60	28.52	192.816	2.35	1.385
9	1848.42	108.58	2.14	50.74	232.3612	1852.22	45.05	2.11	21.35	95.0555	3.80	2.410
10	1860.25	101.24	2.33	43.45	235.8892	1864.22	46.06	1.80	25.59	82.908	4.18	2.198
11	1870.50	86.03	1.64	52.46	141.0892	1872.50	68.36	2.44	28.02	166.7984	2.00	1.258
12	1881.40	35.70	1.21	29.50	43.197	1884.15	63.07	2.24	28.16	141.2768	2.75	0.566
13	1893.41	68.09	2.06	33.05	140.2654	1896.13	33.01	3.15	10.48	103.9815	2.72	1.629
14	1905.69	57.87	2.17	26.67	125.5779	1908.84	41.79	1.91	21.88	79.8189	3.15	1.385
15	1916.89	70.05	1.66	42.20	116.283	1919.26	65.50	2.33	28.11	152.615	2.38	1.069
16	1926.10	40.00	1.07	37.38	42.8	1928.42	81.04	2.49	32.55	201.7896	2.32	0.493
17	1937.33	59.09	1.65	35.81	97.4985	1939.36	81.68	2.60	31.41	212.368	2.03	0.723
18	1947.25	48.56	1.00	48.56	48.56	1948.94	128.76	2.81	52.77	361.8156	1.69	0.377
19	1957.75	66.46	1.50	44.31	99.69	1959.02	170.94	2.44	70.06	417.0936	1.27	0.389
20	1968.42	79.05	2.50	31.62	197.625	1971.48	68.94	3.20	21.54	220.608	3.06	1.146
21	1979.37	76.54	1.36	56.28	104.0944	1981.49	135.15	2.51	53.84	339.2265	2.12	0.566
22	1989.13	87.60	1.20	73	105.12	1991.21	140.05	2.11	66.37	295.5055	2.08	0.625
23	1999.77	107.40	1.55	69.29	166.47	2002.41	96.05	1.80	53.36	172.89	2.71	1.118
24	2010.98	74.98	1.68	44.63	125.9664	2014.05	50.83	2.44	20.83	124.0252	3.07	1.475

Peaks position and their amplitude

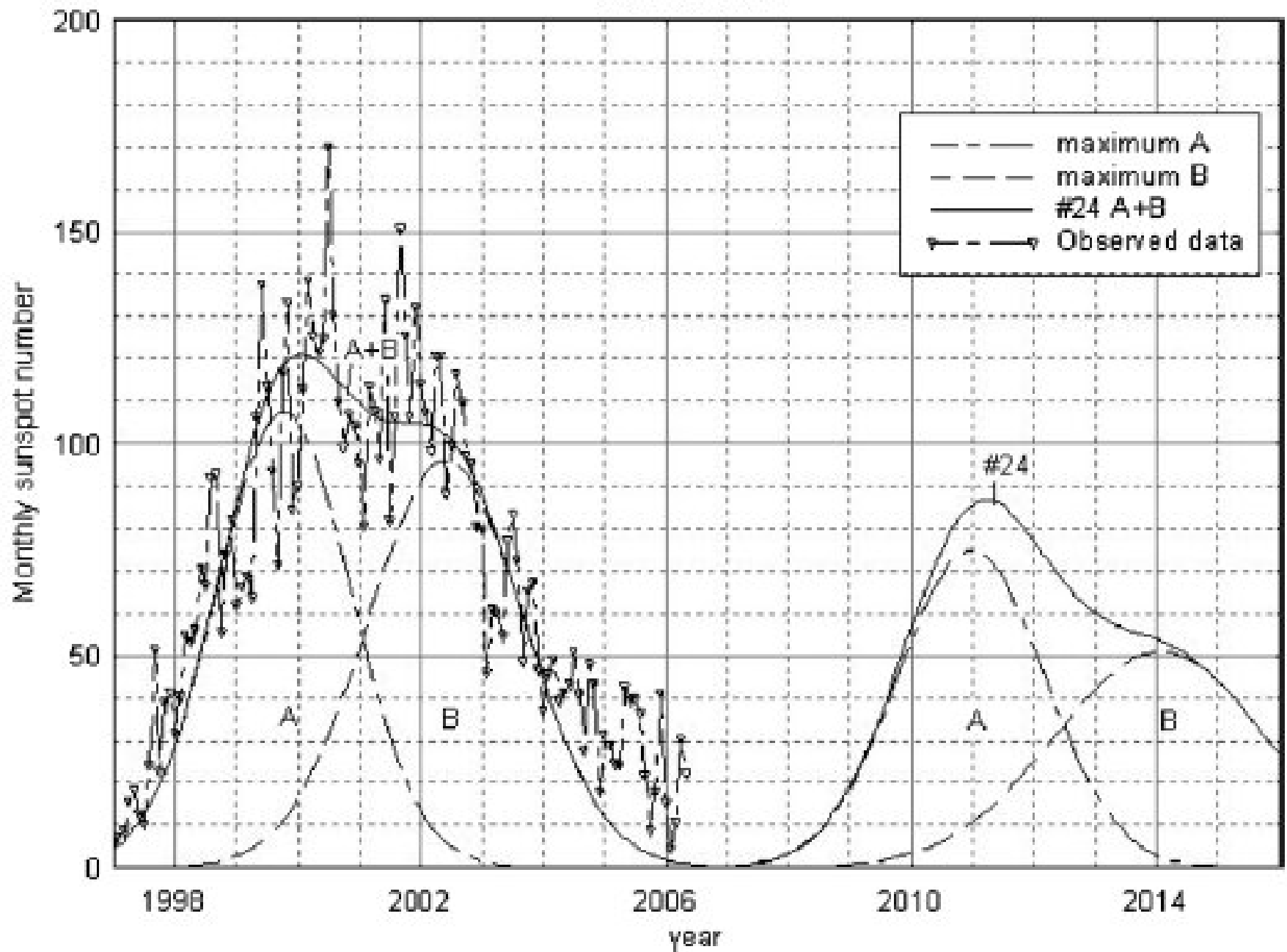




1749 - 2006



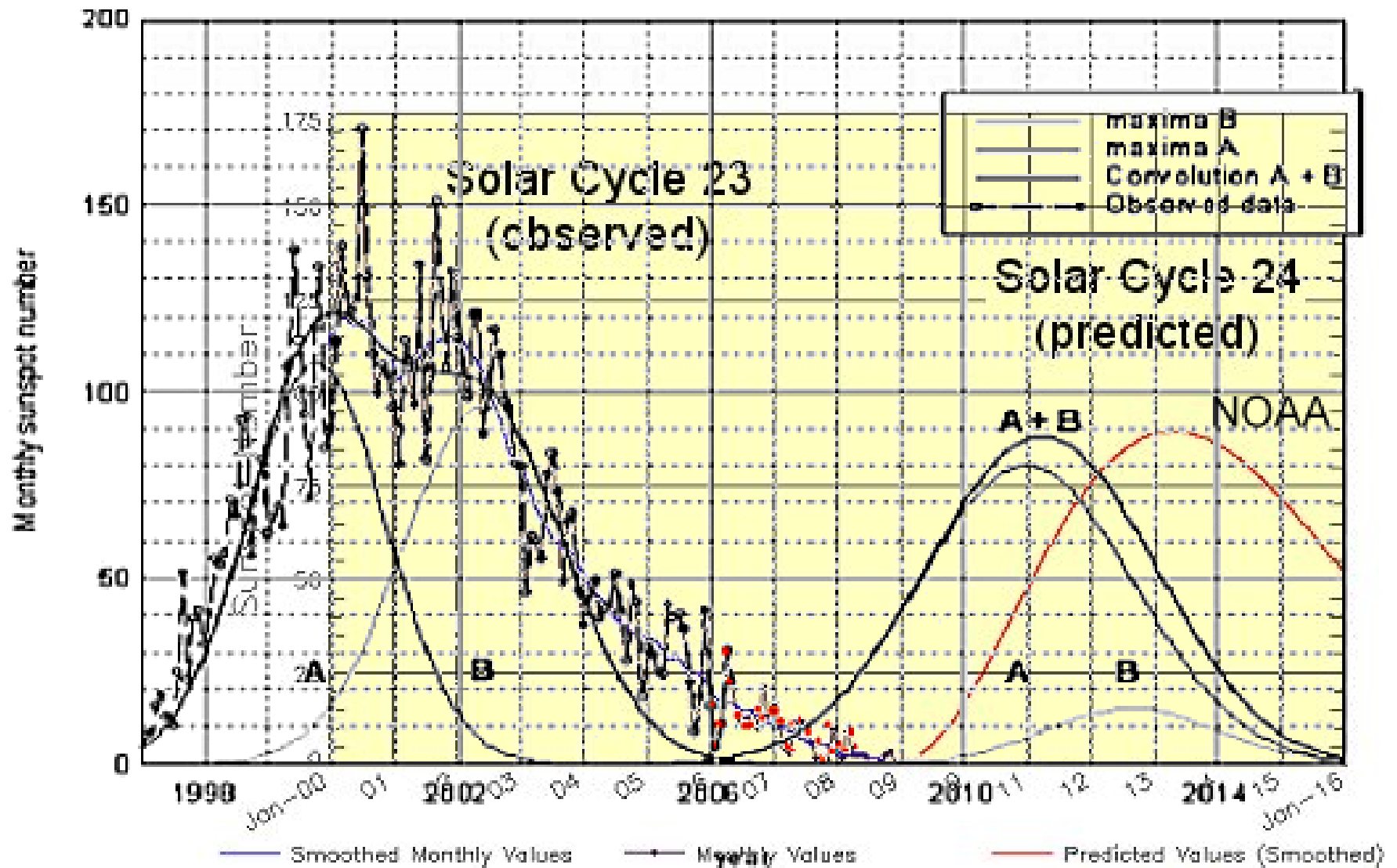
sc #23 + sc#24



(M.V.Rusu, *Advances in Space Research* **40**, 2007)

Forecast of NOAA and ours

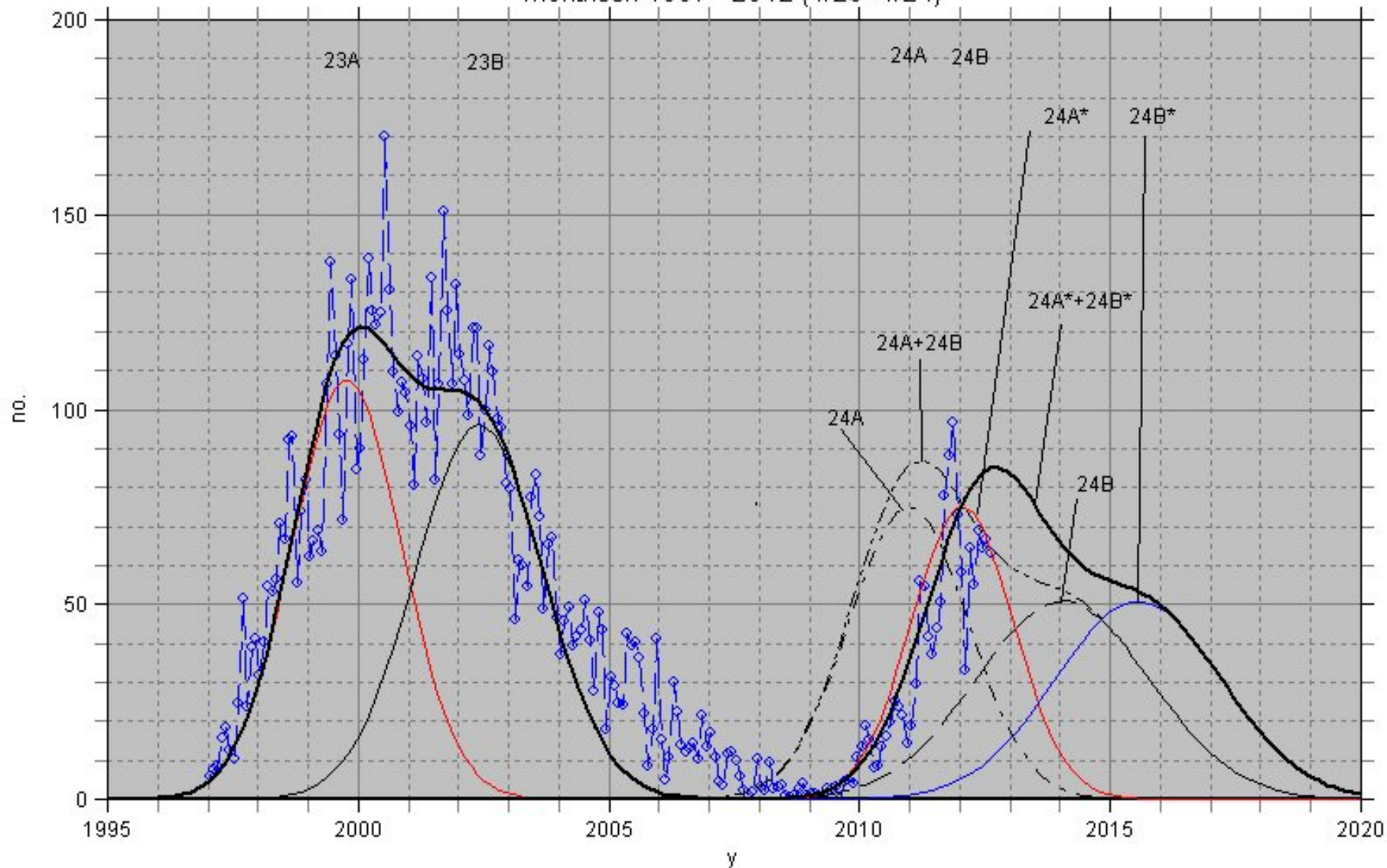
sc #23 + sc#24



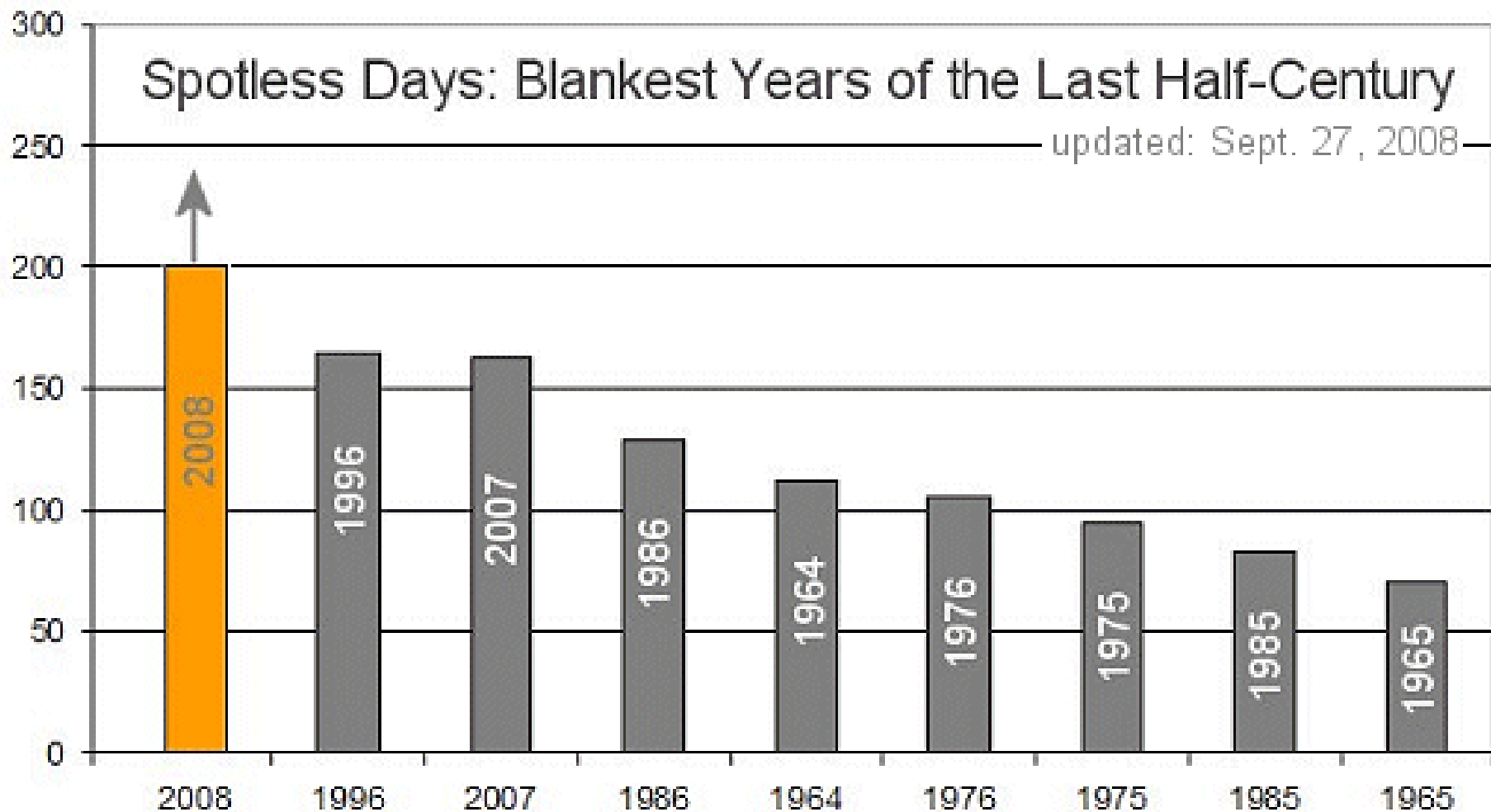
Updated 2009 May 8

NOAA/SWPC Boulder, CO USA

monthssn 1997 - 2012 (#23 - #24)



Prediction and correction for cycle #24

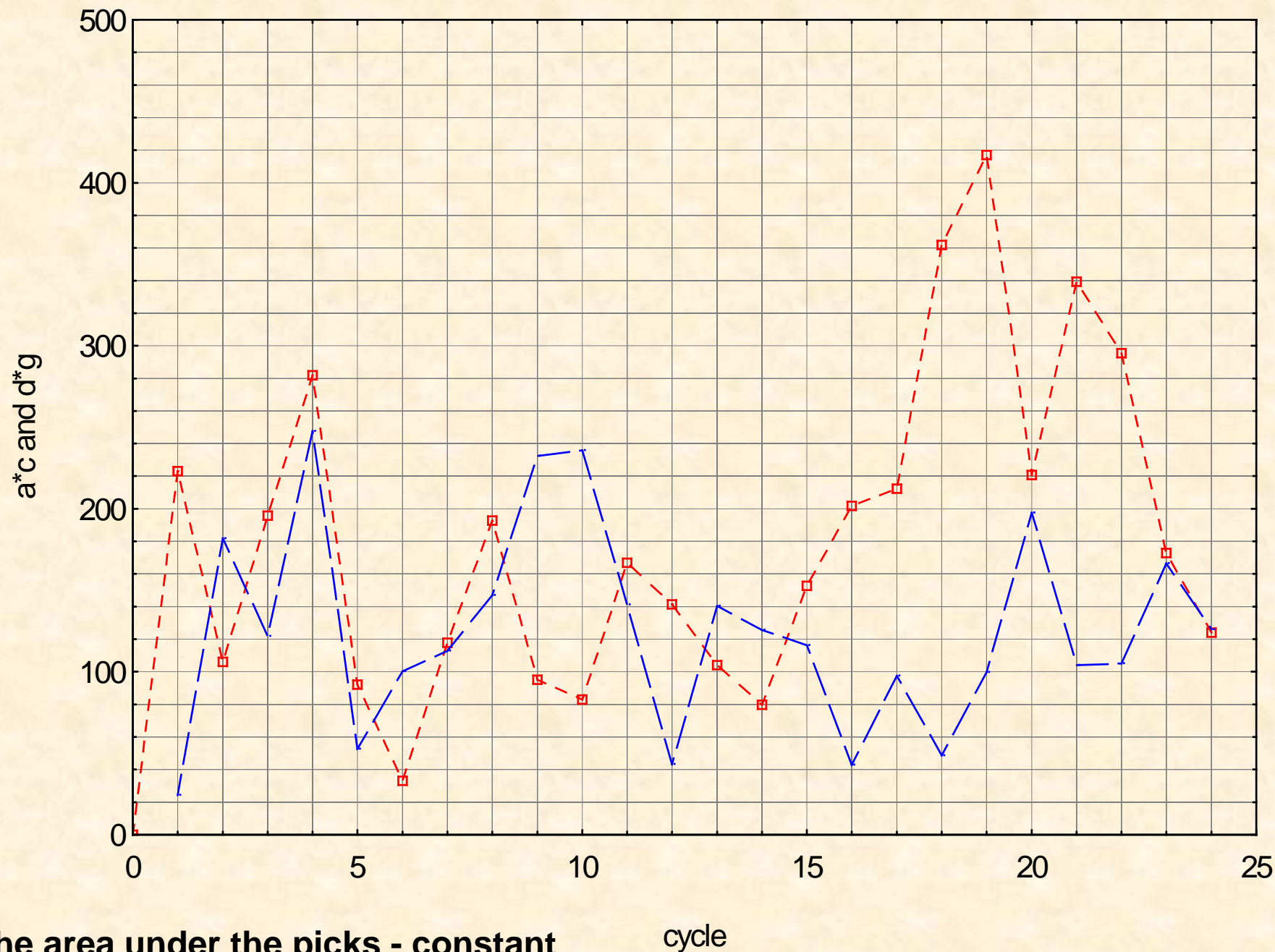


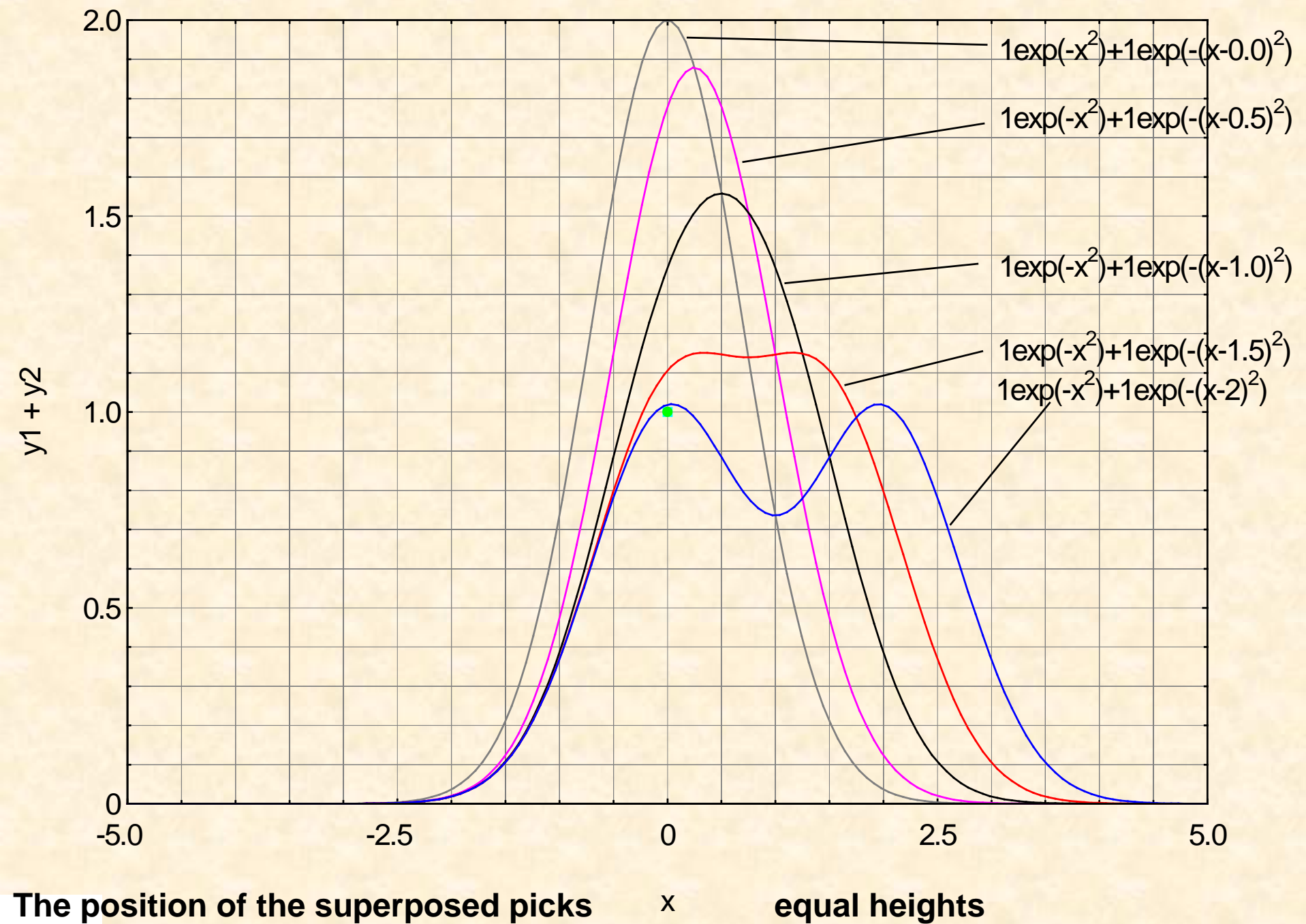
New Solar Cycle Prediction

05.29.2009

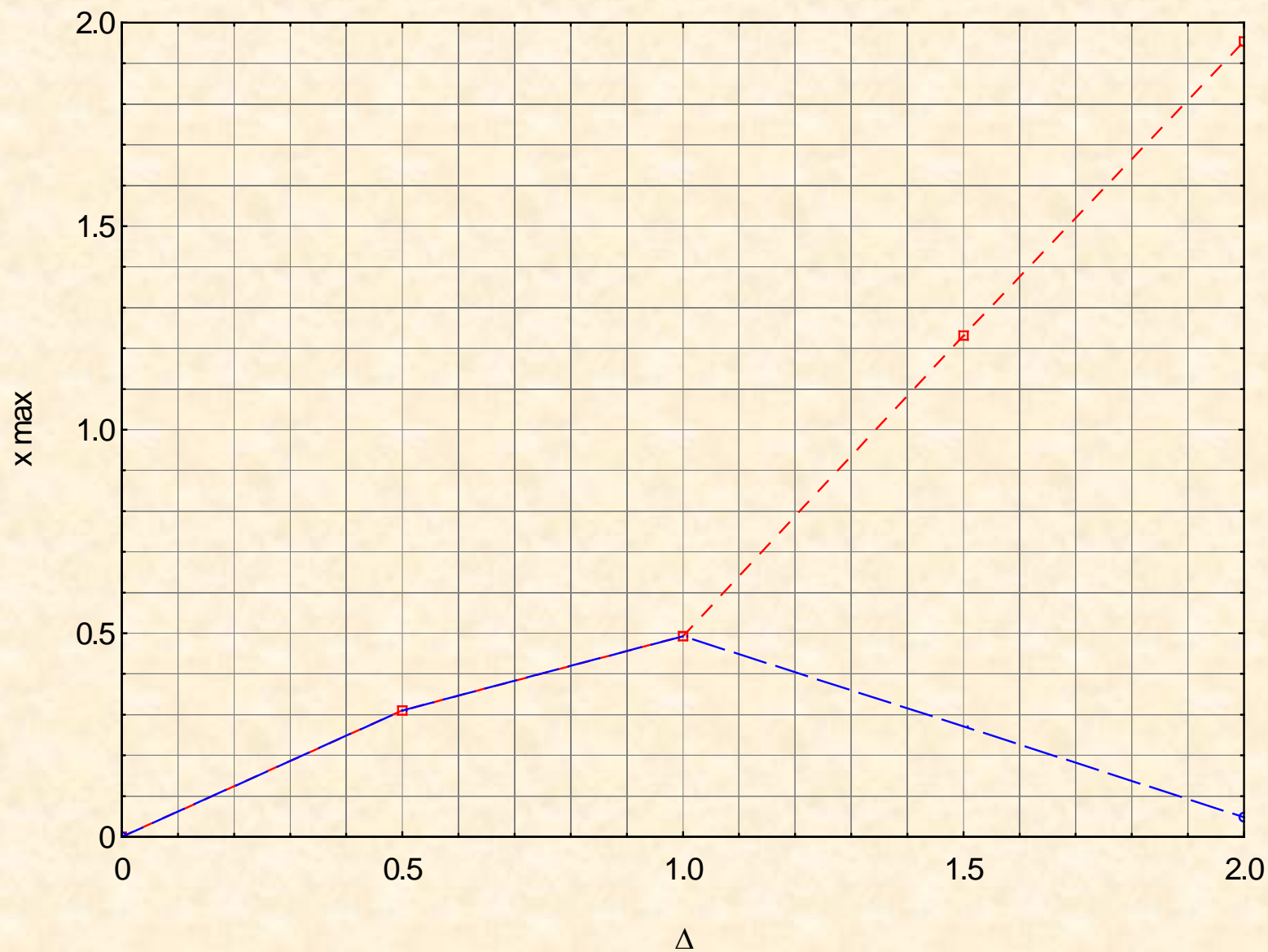
An international panel of experts led by NOAA and sponsored by NASA has released a new prediction for the next solar cycle. Solar Cycle 24 will peak, they say, in May 2013 with a below-average number of sunspots.

a*c for A and d*g for B

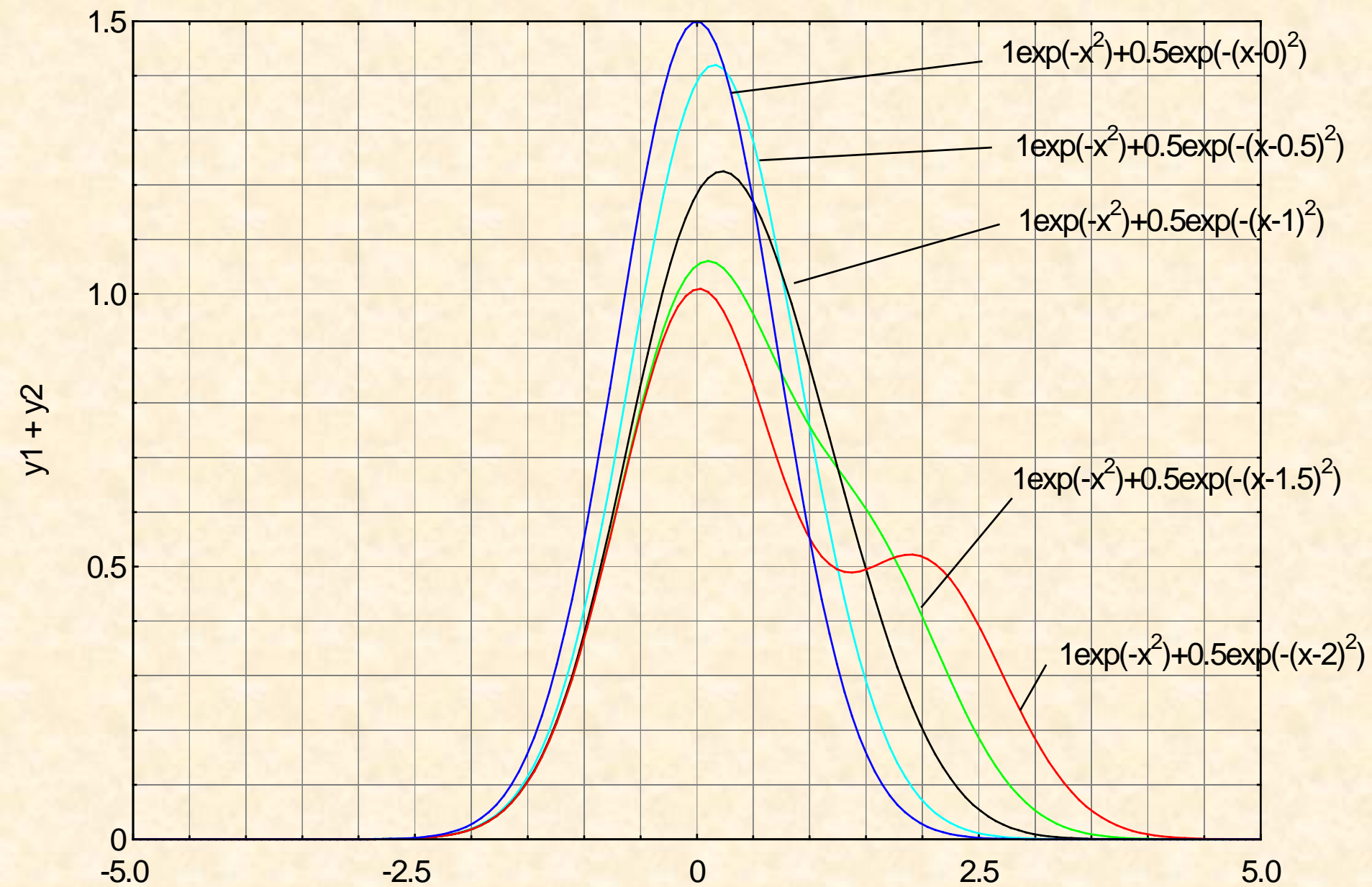




2 gaussians



The position of the superposed picks – the shift

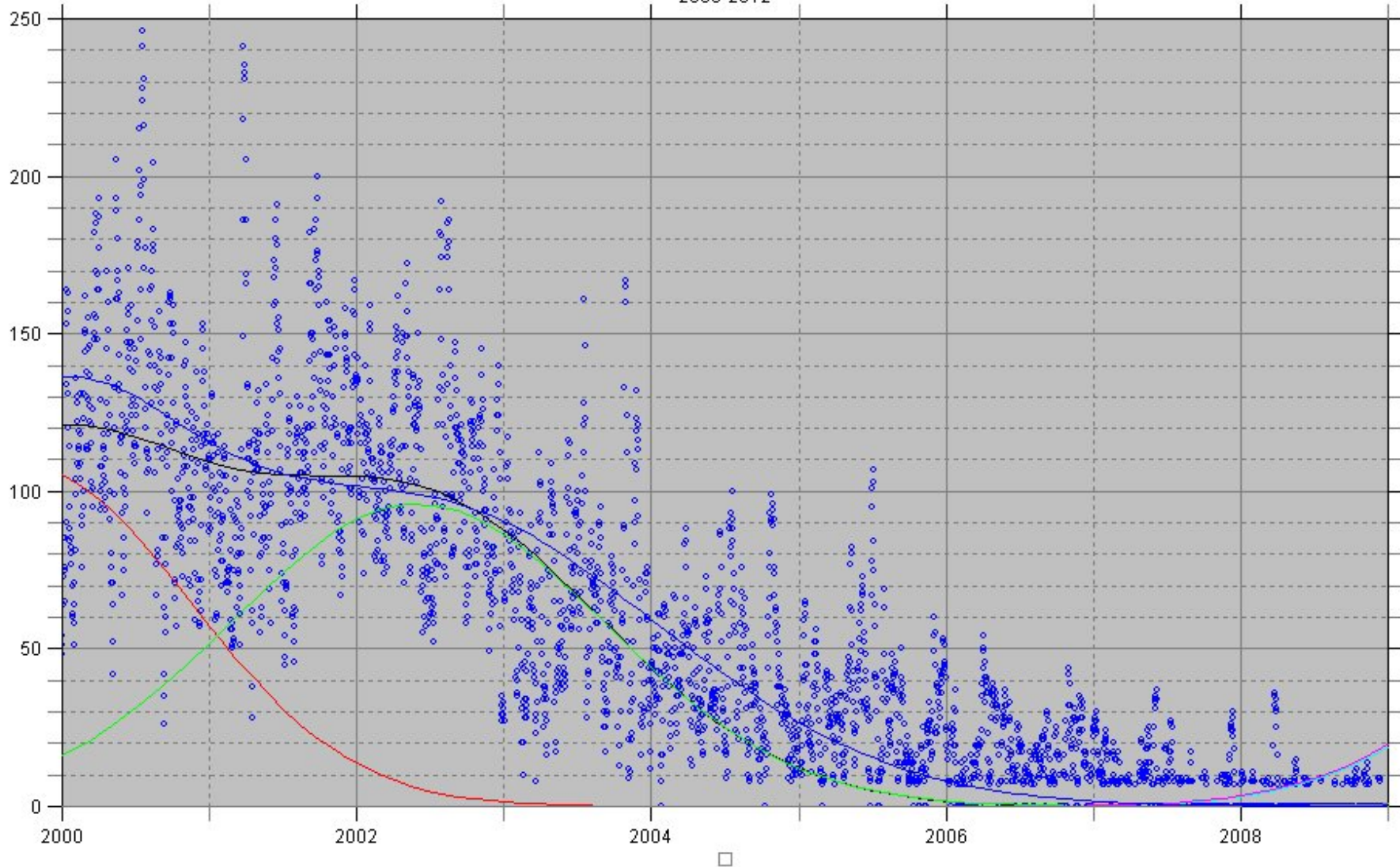


The position of the superposed picks

Δ

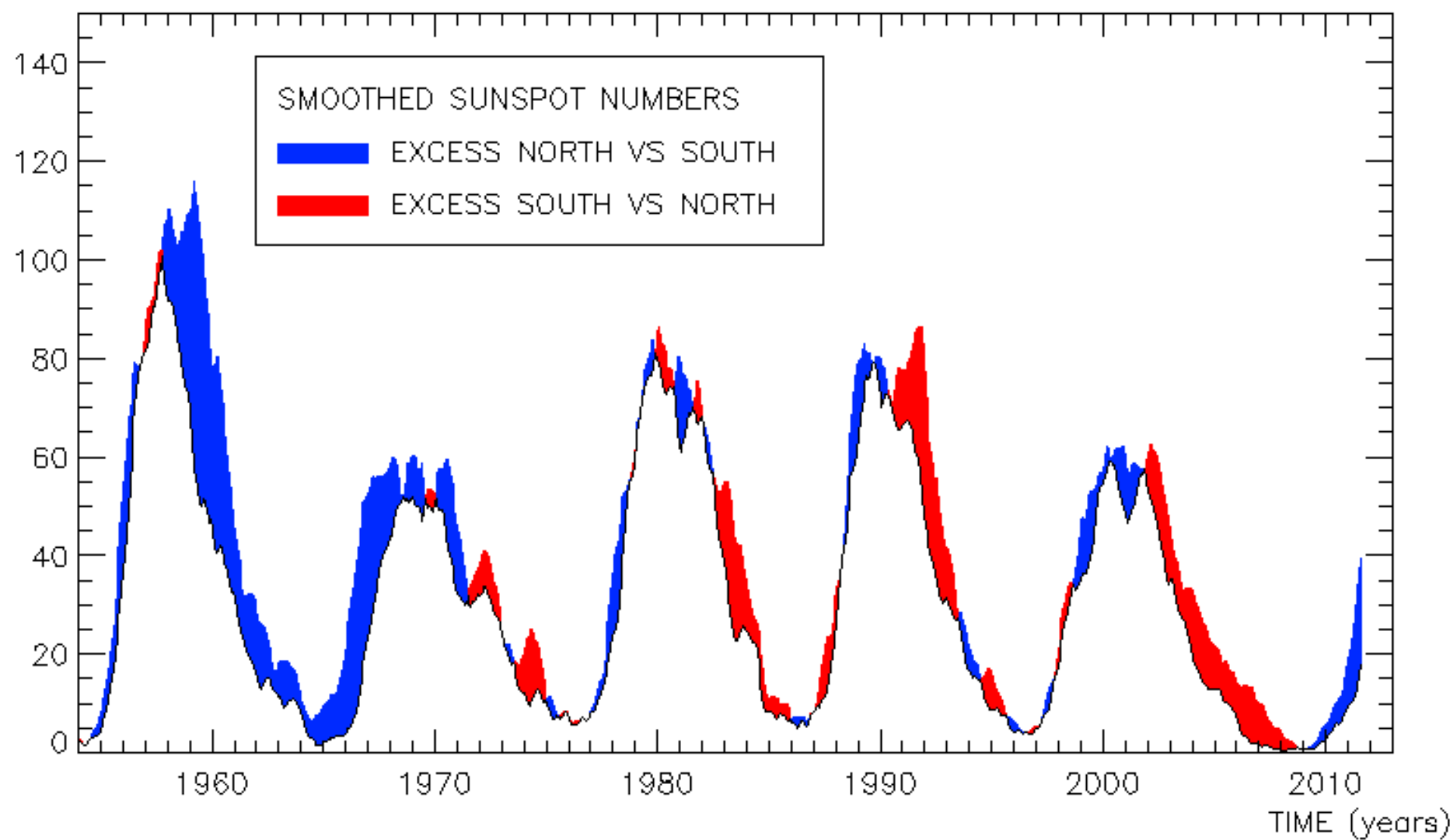
unequal heights

2000-2012



eqn: $107.4\exp(-((x-1999.77)/a)^2) + 96.05\exp(-((x-2002.41)/b)^2)$, R:26.6,
a=1.40629, b=2.28715

The position of the third possible pick



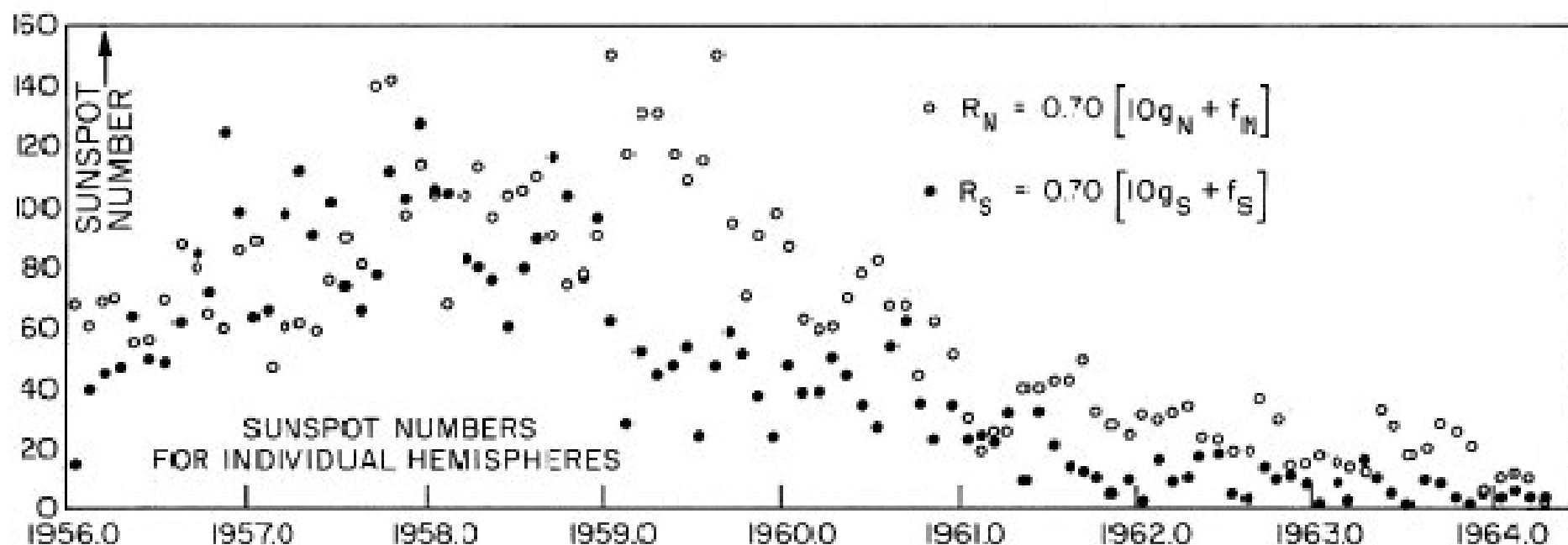


Figure 17

Monthly averages of the sunspot numbers, for the northern and southern hemispheres, derived from Fraunhofer Institut daily maps of the sun during the period 1956-1964. R_N and R_S refer to the sunspot numbers for the northern and southern hemispheres respectively. The g 's refer to the number of groups and the f 's refer to the number of spots.

Conclusions

- New model for sunspot cycle variation having also a physically interpretation besides the statistical one
- Nonlinearly coupled oscillators could be assimilated to a parametric self sustained oscillator
- The periodicity of 11 year becomes more conventional, being difficult to assume the “distance’ between two peak activities
- The “elementary” Gaussian peaks
 - correspond physically to a combination of normal errors distribution
 - are bound to the Green function solution of the nonlinear partial differential equation of the diffusion phenomena that govern surface phenomena of the Sun
- Statistically it is shown that the two peaks are not related to the North- South asymmetry
- The model of two peaks give us analytical function, inferred from the statistically and physically correlated characteristics, that permit forecast of the next cycle, we did for the cycle #24
- Equal area means constant energy in the sunspot systems

Thanks for your attention!



References

- [1] Mircea V. Rusu, Numerical Analysis of the Wolf's Number Time Series; The Asymmetry of its Variation and a Possible Interpretation, The Second International Symposium on Space Climate, Sinaia, Romania, 13-16 Sept. 2006
- [2] Mircea V. Rusu, The Asymmetry of the Solar Cycle: A Result of Non-Linearity, *Advances in Space Research*, **40**, (2007) 1904 – 1911 [1] Mircea V. Rusu, Numerical Analysis of the Wolf's Number Time Series; The Asymmetry of its Variation and a Possible Interpretation, *The Second International Symposium on Space Climate*, Sinaia, Romania, 13-16 Sept. 2006
- [2] Mircea V. Rusu, The Asymmetry of the Solar Cycle: A Result of Non-Linearity, *Advances in Space Research*, **40**, (2007) 1904 - 1911
- Rigozo N.R., Echer E., Vierea L.E.A., Nordemann D.J.R., "Reconstruction of Wolf sunspot numbers on the basis of spectral characteristics and estimates of associated radio flux and solar wind parameters for the last millennium", 2001: *Solar Physics*, **203**: 179-191.
- Hoyt, D. V. and Schatten, K. H.: "A new look at Wolf sunspots numbers in the late 1700's", *Solar Physics*, 1992: **138**: 387-397.
- Sello, S., "Wavelet analysis of solar activity", Topic Note Nr. UGS/12012000 (V1.1).
- Mursula, K., Usoskin, I. G. and Kovaltsov, G. A. "Persistent 22-year cycle in sunspot activity: evidence for a relic solar magnetic field", 2001: *Solar Physics* **198**: 51-56.
- Weiss, N.O. and Tobias, S.M. "Physical causes of solar activity", 2000: *Space Science Reviews* **94**: 99-112.
- Lu, E. T., Hamilton, R. J., Tiernan, J. M. and Bromund, K. R. "Solar flares and avalanches in driven dissipative systems", 1993: *The Astrophysical Journal*, **412**: 841-852.
- Nagata, S., Hara, H., Kano, R., Kobayashi, K., Sakao, T., Shimizu, T., Tsuneta, S., Yoshida, T. and Gurman, J. B.: "Spatial and temporal properties of hot and cool coronal loops", 2003: *The Astrophysical Journal*, **590**: 1095-1110.
- Burlaga, L. F., Wang, C., Richardson, J. D. and Ness, N. F.: "Evolution of magnetic fields in corotating interaction regions from 1 to 95 AU: order to chaos", 2003: *The Astrophysical Journal*, **590**: 554-566.

- Carrol B.W., Ostlie D.A., *Modern Astrophysics*, Addison-Wesley Publ. 1996
- Komitov, B. and Bonev, B., "Amplitude variations of the 11 year cycle and the current solar maximum 23", *The Astrophysical Journal*, 2001: 554: L119-L122, June 10.
- Allen C.W., *Astrophysical Quantities*, Univ. of London, 1964
- Zirin H., *Astrophysics of the Sun*, Cambridge Univ. Press, 1988
- Verma V. K., "Periodic Variation of the North-South Asymmetry of Solar Activity Phenomena", *J. Astrophys. Astr.* (2000) **21**, 173-176
- Verma, V. K. 1987, *Solar Phys.*, **114**, 185, 1992, *ASP Conf. Series*, **27**, 429. 1993, *Astrophys. J.*, **403**, 797, 2000, *Solar Phys.*, **194**, 87.
- Roy, J. R. 1977, *Solar Phys.*, **52**, 53.
- Baranovski A., Clette F., Nollau V., Nonlinear solar cycles forecasting: theory and perspectives,
- Kilcik A., Andreson C.N., Rozelot J.P., Ye H., Sugihara G., Osguc A. Nonlinear Prediction of Solar Cycle 24,
- Maris G., Popescu M.D., Besliu D., 2004 Symposium 223, Cambridge University Press, Cambridge, 127,
- Maris G., Oncica A. 2006, *Sun Geosph.*, 1, 8
- Poligiannakis J.M., Moussas X., A nonlinear RLC Solar Cycle Model, *Solar Physics* **163**, 193-203, 1996
- Ahlumwalia, The forecast for solar cycle 2 activity: A progress report, *Proceedings of ICRC 2001*: 3359
- Hiremath K.M., Prediction of solar cycle 24 and beyond, 2008,

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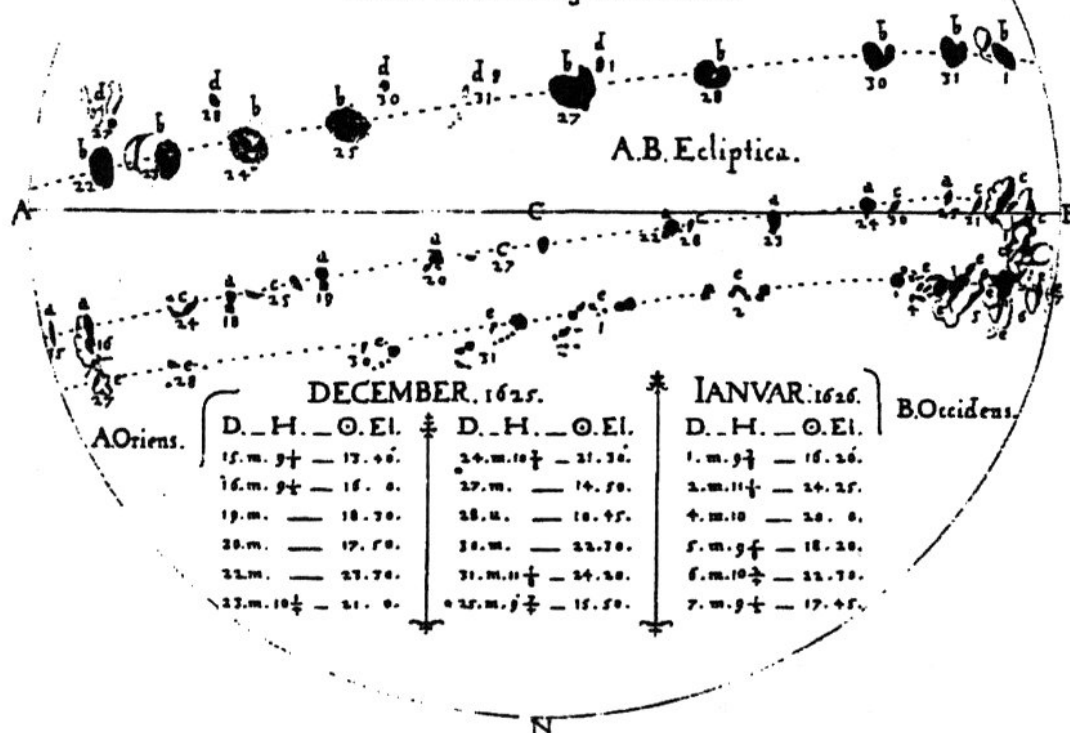
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In Collegio Romano Societatis

6. Decembris.
1. Ianuarij. 1626.

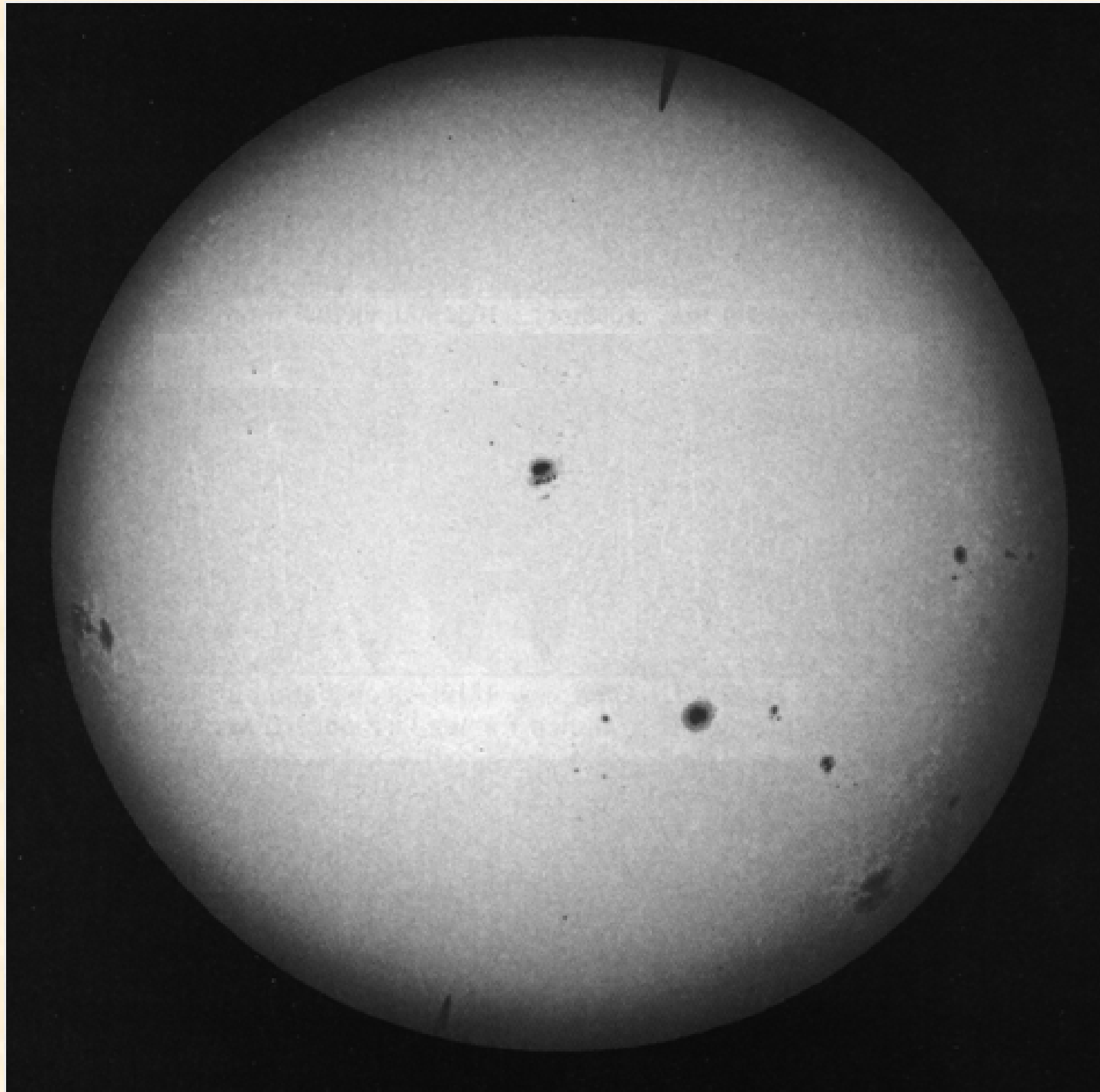


10
9
7
f

Curfus Macularum à 15. Dec: anni 1625, ad 7. Ian: anni 1626.
similis alijs alioru annorum.



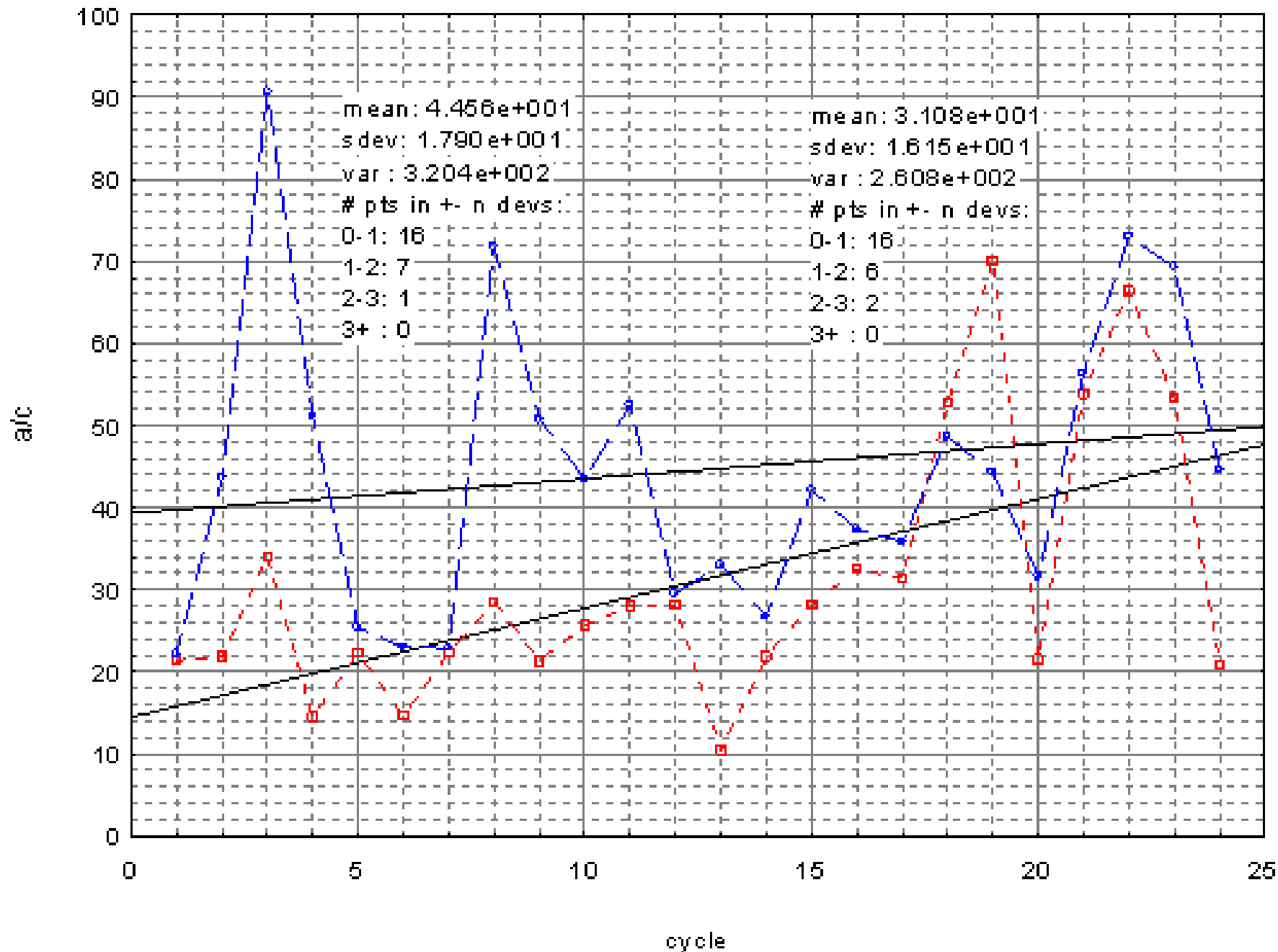
Early sunspot observations. This figure, from Scheiner's *Rosa Ursina*, completed in 1630, shows several sunspots on successive days, as they are carried round by the Sun's rotation. The distinction between the umbra and penumbra of a sunspot had already been recognized. (Courtesy of the RAS.)



a/c for A and B

$$y = +0.420x^1 + 39.3, R: 18.0, \text{max dev: } 50.0$$

$$y = +1.33x^1 + 14.5, R: 13.4, \text{max dev: } 30.3$$

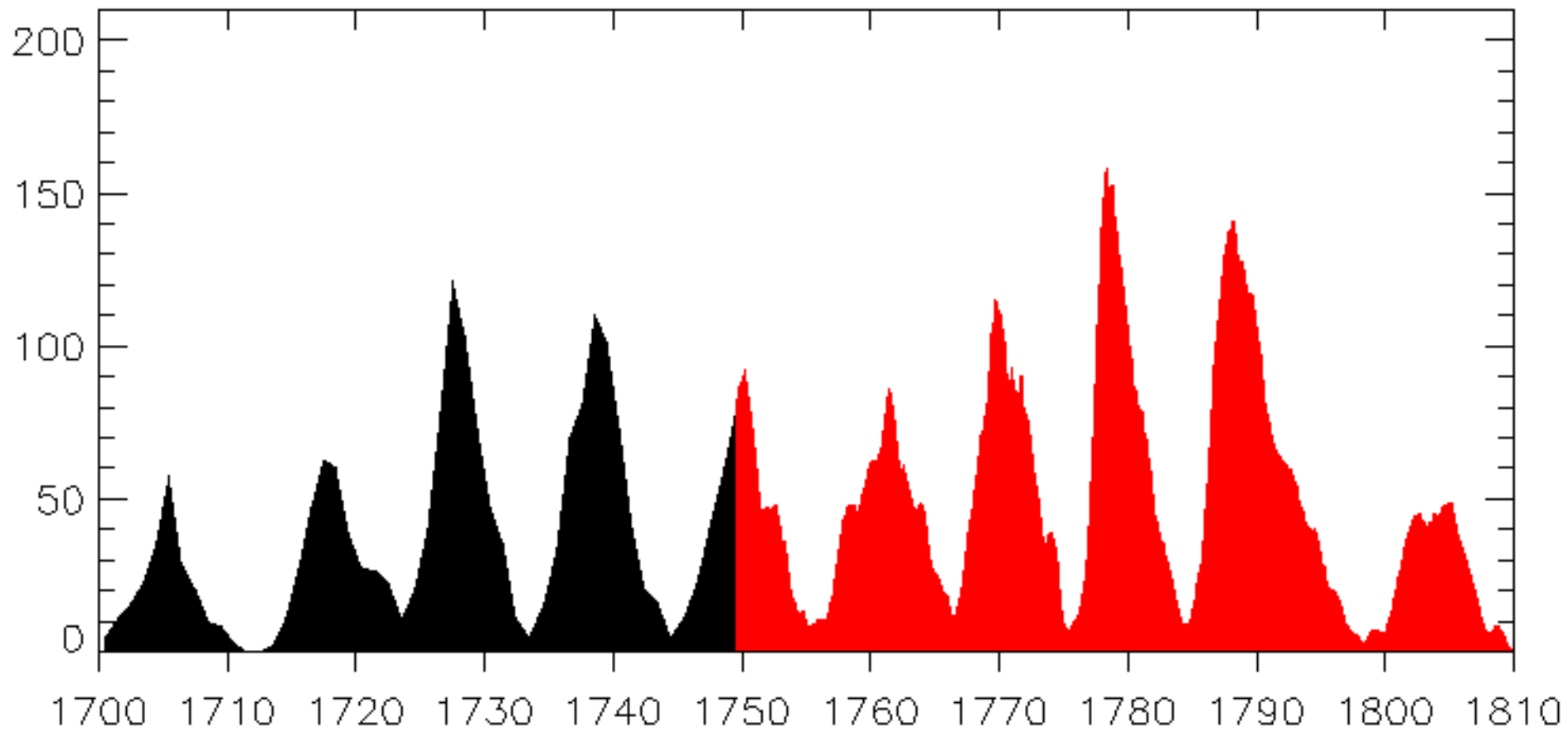


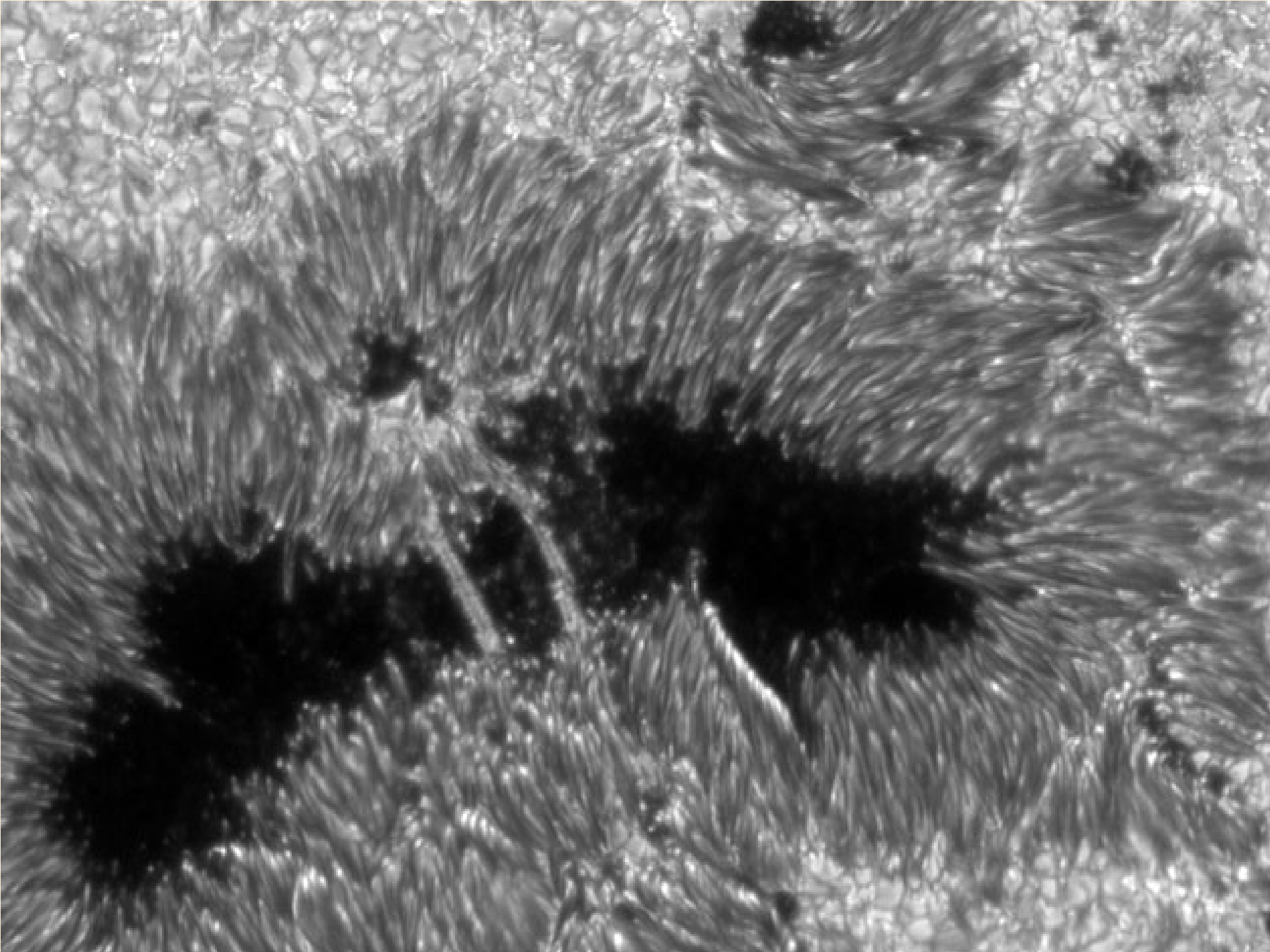
No.	A				B				Δ	R
cycle	b poz 1	a ampl 1	c width 1	a/c	f poz 2	d ampl 2	g width 2	d/g	f - b	a/d
1	1759.03	23.21	1.05	22.10	1761.78	69.07	3.23	21.38	2.75	0.336
2	1769.67	89.10	2.04	43.67	1772.37	48.19	2.20	21.90	2.70	1.849
3	1778.35	105.07	1.16	90.58	1780.22	81.59	2.40	34.00	1.87	1.288
4	1788.03	112.51	2.20	51.14	1791.40	64.09	4.40	14.56	3.37	1.755
5	1802.05	36.54	1.44	25.37	1804.89	45.34	2.03	22.33	2.84	0.806
6	1816.20	48.00	2.09	22.97	1819.50	22.03	1.50	14.69	3.30	2.179
7	1828.04	50.79	2.22	22.88	1830.00	51.23	2.30	22.27	1.96	0.991
8	1836.88	102.71	1.43	71.82	1839.23	74.16	2.60	28.52	2.35	1.385
9	1848.42	108.58	2.14	50.74	1852.22	45.05	2.11	21.35	3.80	2.410
10	1860.25	101.24	2.33	43.45	1864.22	46.06	1.80	25.59	4.18	2.198
11	1870.50	86.03	1.64	52.46	1872.50	68.36	2.44	28.02	2.00	1.258
12	1881.40	35.70	1.21	29.50	1884.15	63.07	2.24	28.16	2.75	0.566
13	1893.41	68.09	2.06	33.05	1896.13	33.01	3.15	10.48	2.72	1.629
14	1905.69	57.87	2.17	26.67	1908.84	41.79	1.91	21.88	3.15	1.385
15	1916.89	70.05	1.66	42.20	1919.26	65.50	2.33	28.11	2.38	1.069
16	1926.10	40.00	1.07	37.38	1928.42	81.04	2.49	32.55	2.32	0.493
17	1937.33	59.09	1.65	35.81	1939.36	81.68	2.60	31.41	2.03	0.723
18	1947.25	48.56	1.00	48.56	1948.94	128.76	2.81	52.77	1.69	0.377
19	1957.75	66.46	1.50	44.31	1959.02	170.94	2.44	70.06	1.27	0.389
20	1968.42	79.05	2.50	31.62	1971.48	68.94	3.20	21.54	3.06	1.146
21	1979.37	76.54	1.36	56.28	1981.49	135.15	2.51	53.84	2.12	0.566
22	1989.13	87.60	1.20	73	1991.21	140.05	2.11	66.37	2.08	0.625
23	1999.77	107.40	1.55	69.29	2002.41	96.05	1.80	53.36	2.71	1.118
24	2010.98	74.98	1.68	44.63	2014.05	50.83	2.44	20.83	3.07	1.475

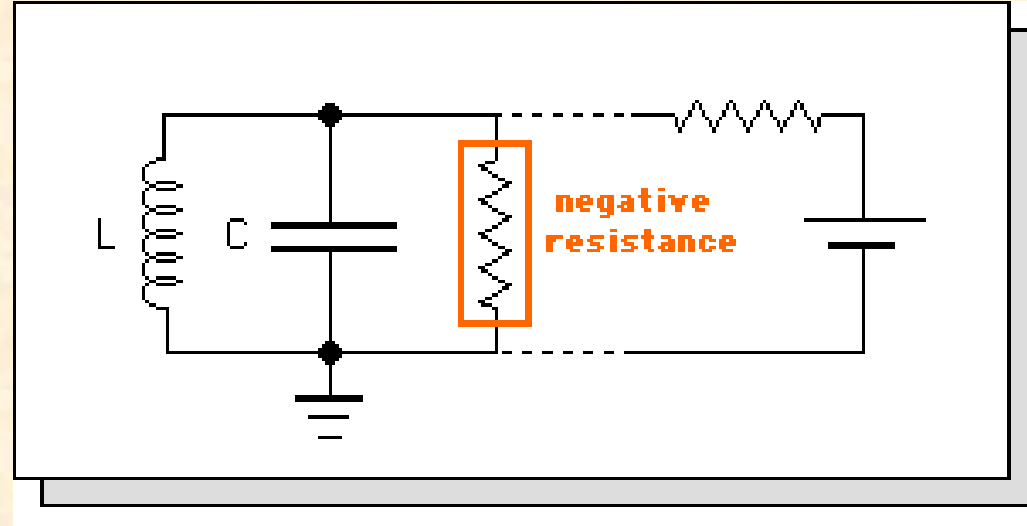
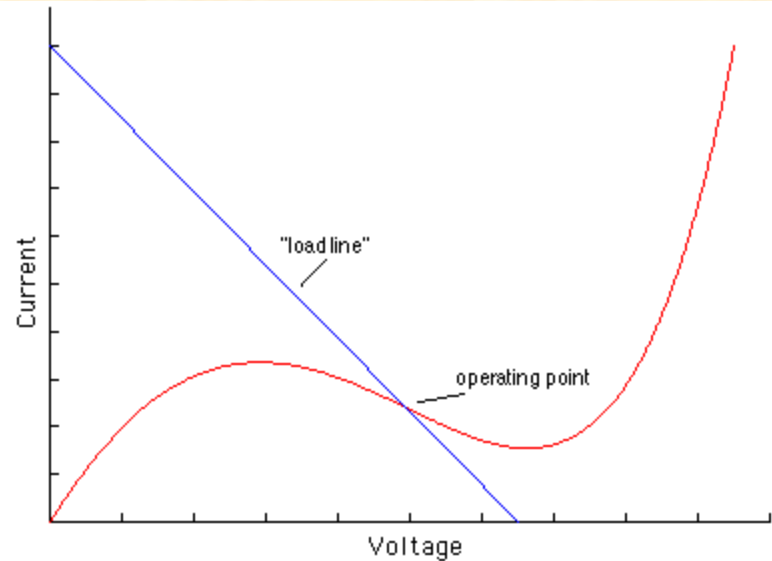
Abstract

- In an early paper we analyzed the asymmetry of the sunspot cycles from the point of view of its non-linearity. There are many attempts to predict the features of the future evolution of the solar activity which are based on different models for fitting the shape of Wolf number evolution. None of them are fully satisfying and the recent delay of the solar cycle #24 is one example. Unfortunately the majority of the predictions are based on the numerical analysis, and just a few attempts to base the forecast on a physical basis. We try to go closer to a solution of the prediction problem, using some physical evidences we found presented in the early article and now we made a critical analysis of our prediction based on the accumulated data that from 2007.

The pattern of the time series of wolf numbers for a log period







$$\frac{d^2}{dt^2} v(t) - \frac{d}{dt} [\alpha v(t) - \beta v^3(t)] + \omega_0^2 v(t) = 0$$

$$\frac{1}{2}(\omega_0^2 - \omega^2) V(t) - i\omega \dot{V}(t) + i\omega \frac{1}{2} [\alpha - \beta |V(t)|^2] V(t) = i \frac{\omega^2 V_0}{2}$$

B. van der Pol, *Radio Rev.* 1, 704-754, 1920 and B. van der Pol, *Phil. Mag.* 3, 65, 1927

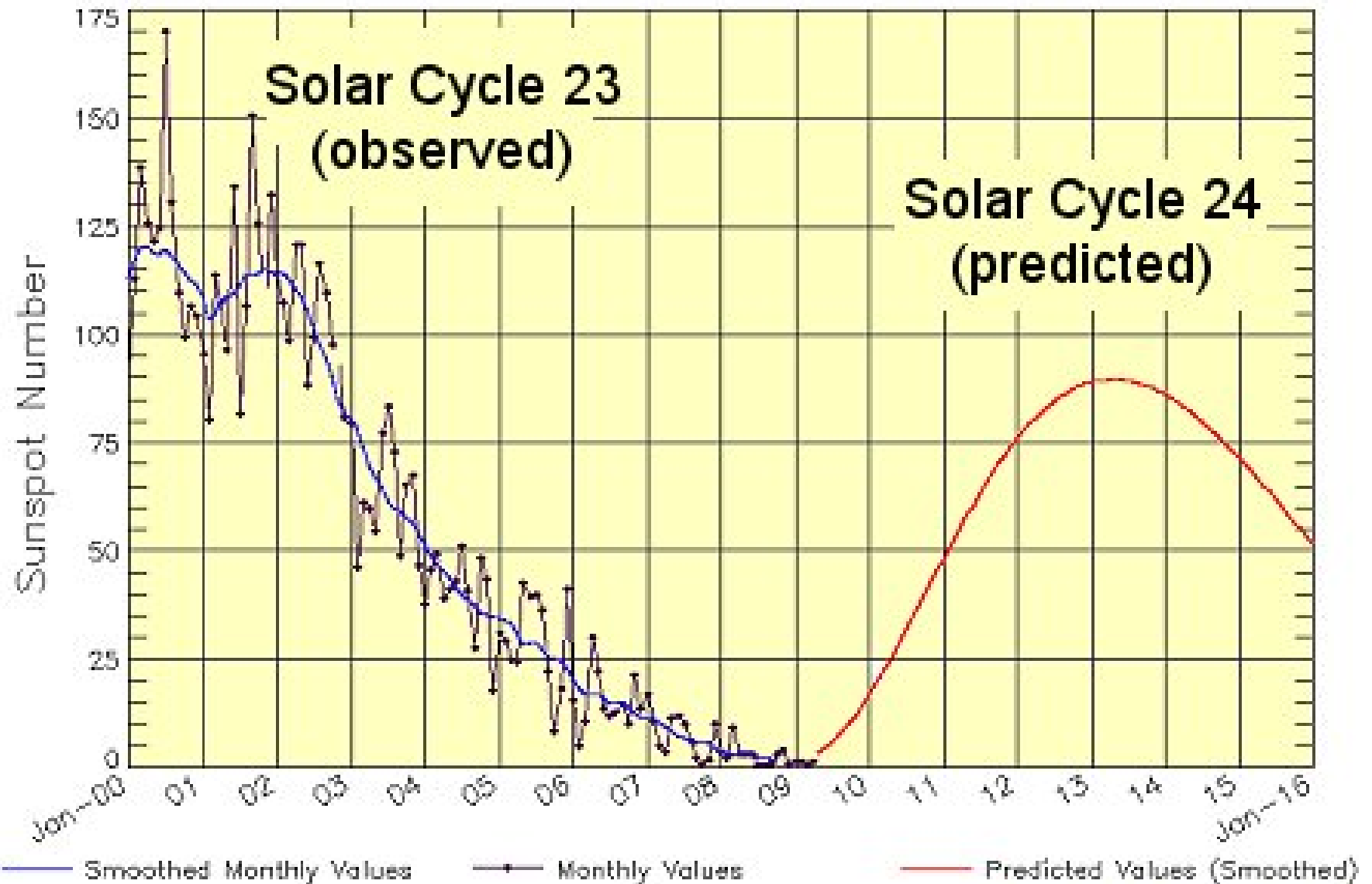
$$y'' - \mu(1 - y^2)y' + y = 0.$$

Parametric oscillator

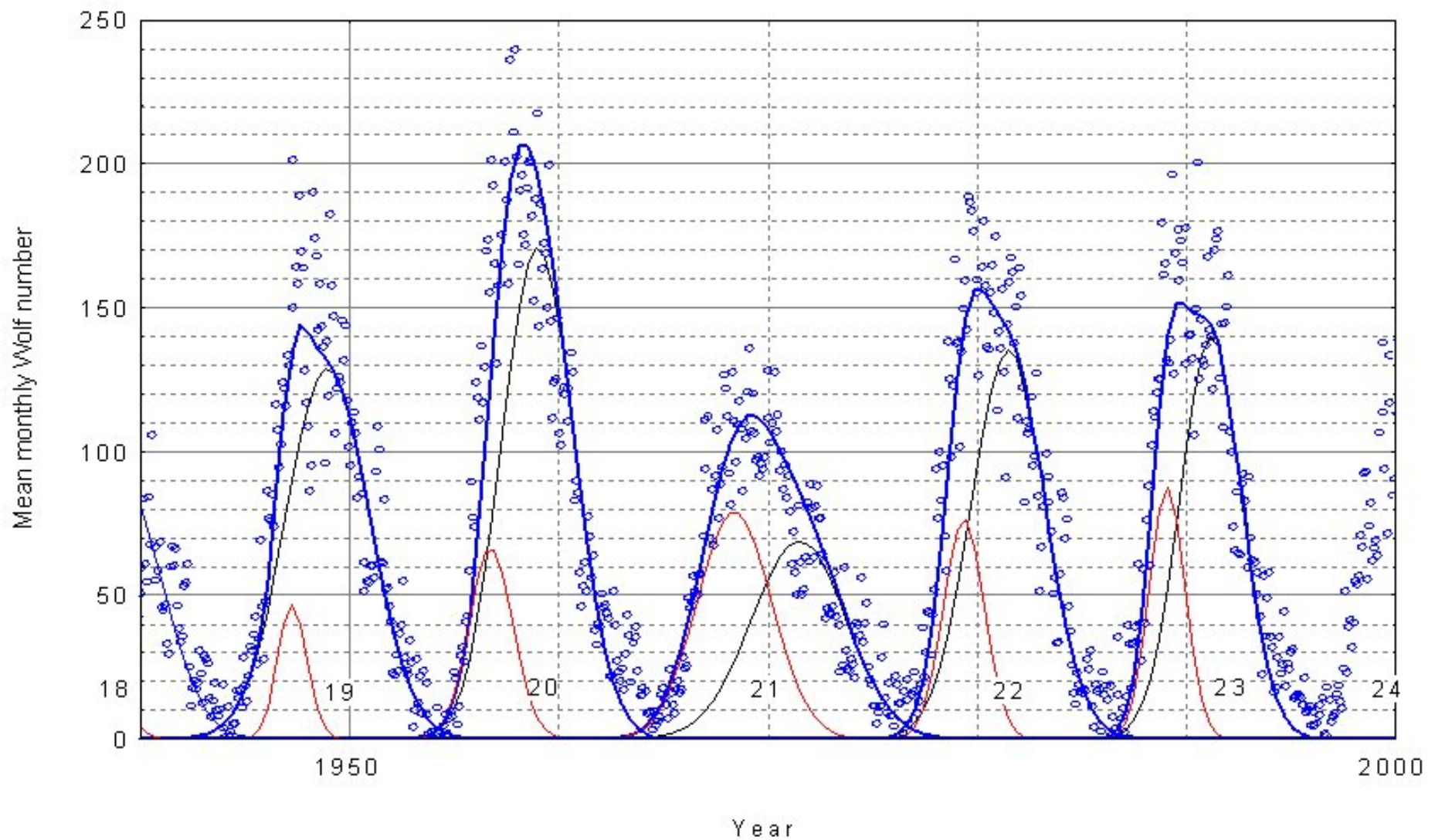
$$\frac{d^2 x}{dt^2} + \beta(t) \frac{dx}{dt} + \omega^2(t)x = 0$$

ISES Solar Cycle Sunspot Number Progression

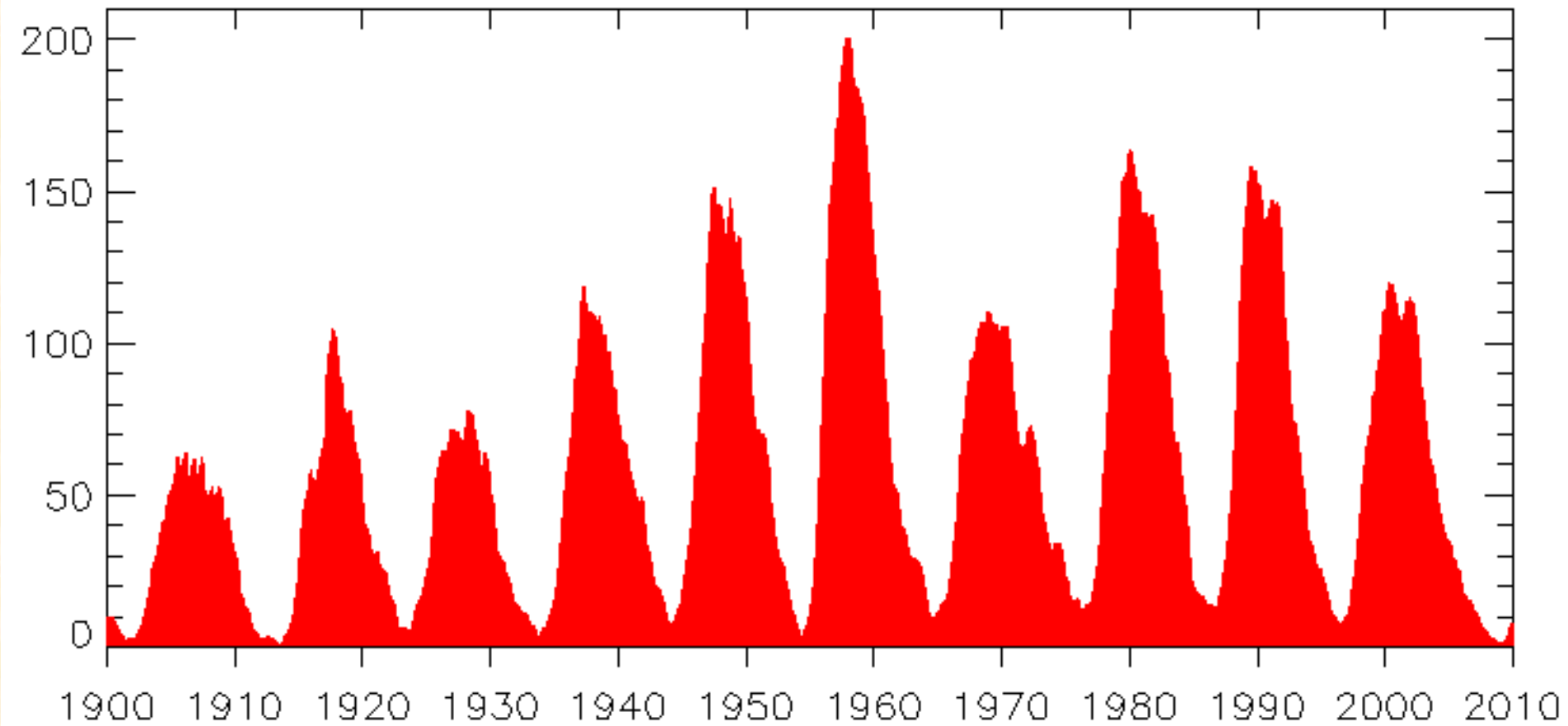
Data Through Apr 09



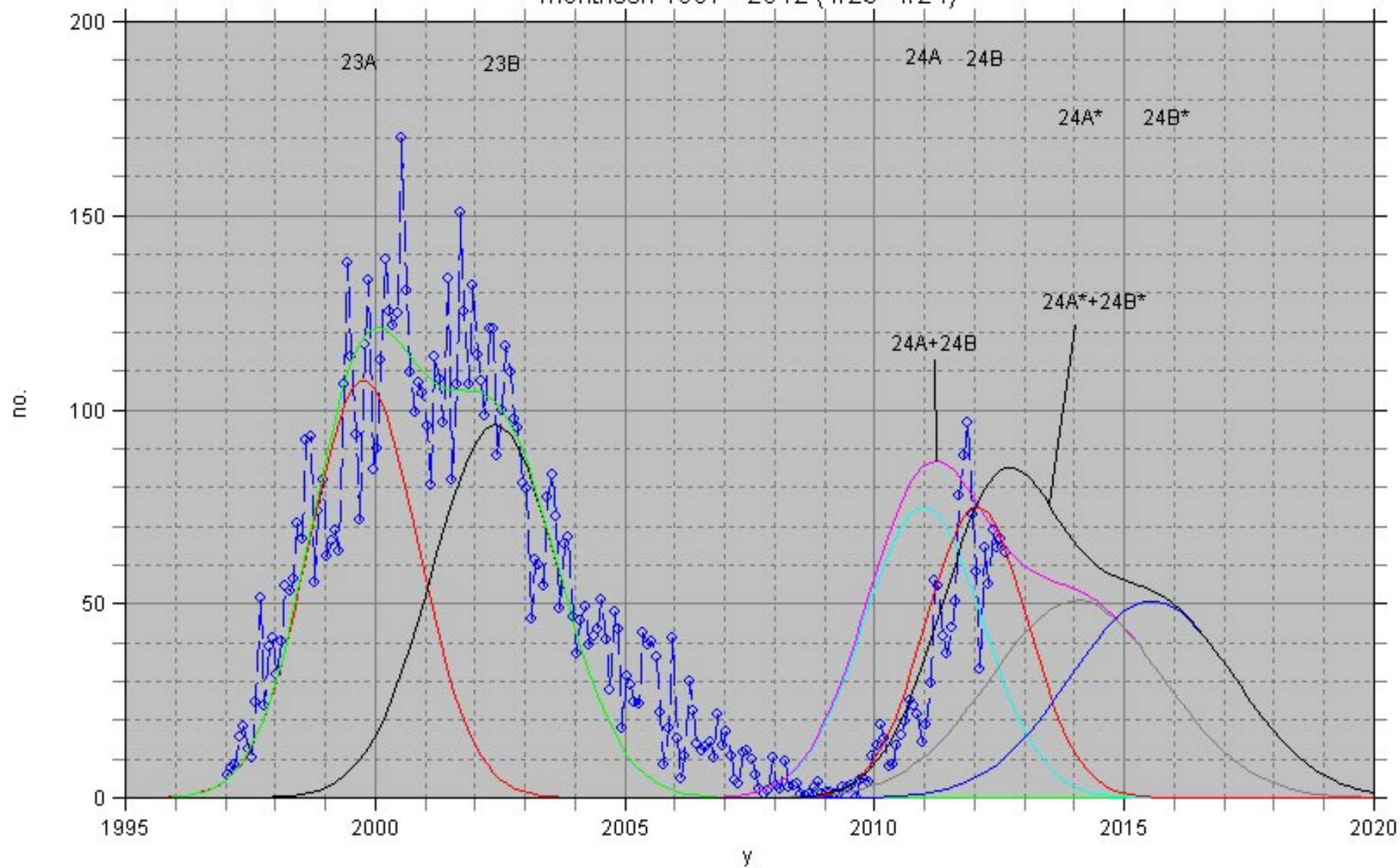
Full set of monthly mean Wolf number



The pattern of the time series of wolf numbers for a log period



monthssn 1997 - 2012 (#23 - #24)



$$\Delta=f-b$$

